

# Fluctuation of Extragalactic Reference Frame Due to Gravitational Lensing

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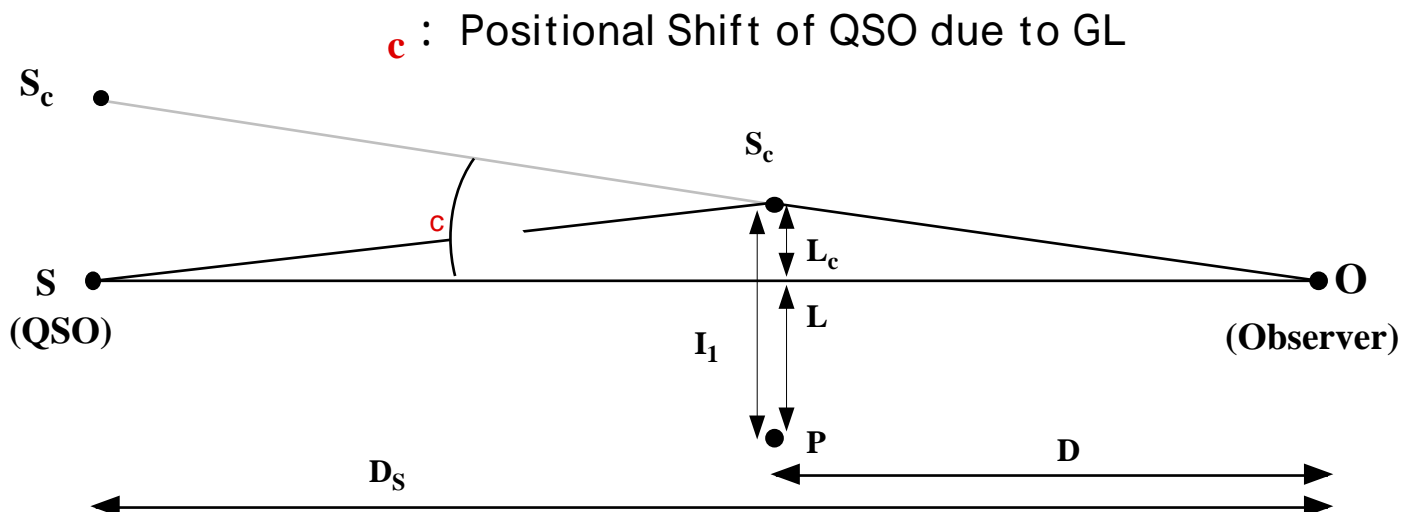
NAO

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Nagano NCT

Observation of QSOs through our galaxy

Observed position of QSOs are shifted,  
apparent proper motion of QSOs are induced  
due to gravitational lensing of the matters in our galaxy.



$$c = \frac{mL}{L^2 + 2mD} : \text{Angular Shift of Light Centroid}$$

$$\text{where } m = \frac{4GM}{c^2}$$

# Our Model

1. Disk Star (  $1 M_{\text{sun}}$  , Exponential Disk model)

$$\rho_D(x, y, z) = \frac{\Sigma_D}{2z_D} \exp \left[ - \left( \frac{\sqrt{x^2 + y^2} - R_0}{R_D} \right) - \left( \frac{z}{z_D} \right) \right]$$

$\Sigma_0 = 46 M_{\text{sun}}/\text{pc}^2$  : Column Density of Disk Stars

$R_0 = 8.5 \text{ kpc}$  : Distance Galactic Center

$z_h = 300 \text{ pc}$  : Disk Thickness

$v_0 = 220 \text{ km/s}$  : flat rotation curve

2. MACHO (  $0.1 M_{\text{sun}}$  , Isothermal model)  $\rho_M = \rho_0 \frac{a^2 + R_0^2}{a^2 + R^2}$

$a$  : Core Radius of MACHO , 2000 pc

$\rho_0$  : Local Density of Dark Matter,  $8 \times 10^{-3} M_{\text{sun}}/\text{pc}^3$

$v = 180 \text{ km/s}$  : Random Direction

We cannot see QSO beyond bulge.

## Expectation Values of QSO positional Shift

Root Mean Square of positional shift  $\theta_c$

$$\langle \theta_c^2 \rangle = \int n(m, \mathbf{r}) \{ \theta_c(m, L) \}^2 dm dV$$

$$= \sum_{i=D, M} \int \int 2\pi \rho_i(D) \frac{L^3 m_i}{(L^2 + 2m_i D)^2} dL dD$$

Mass  $\times$  Density

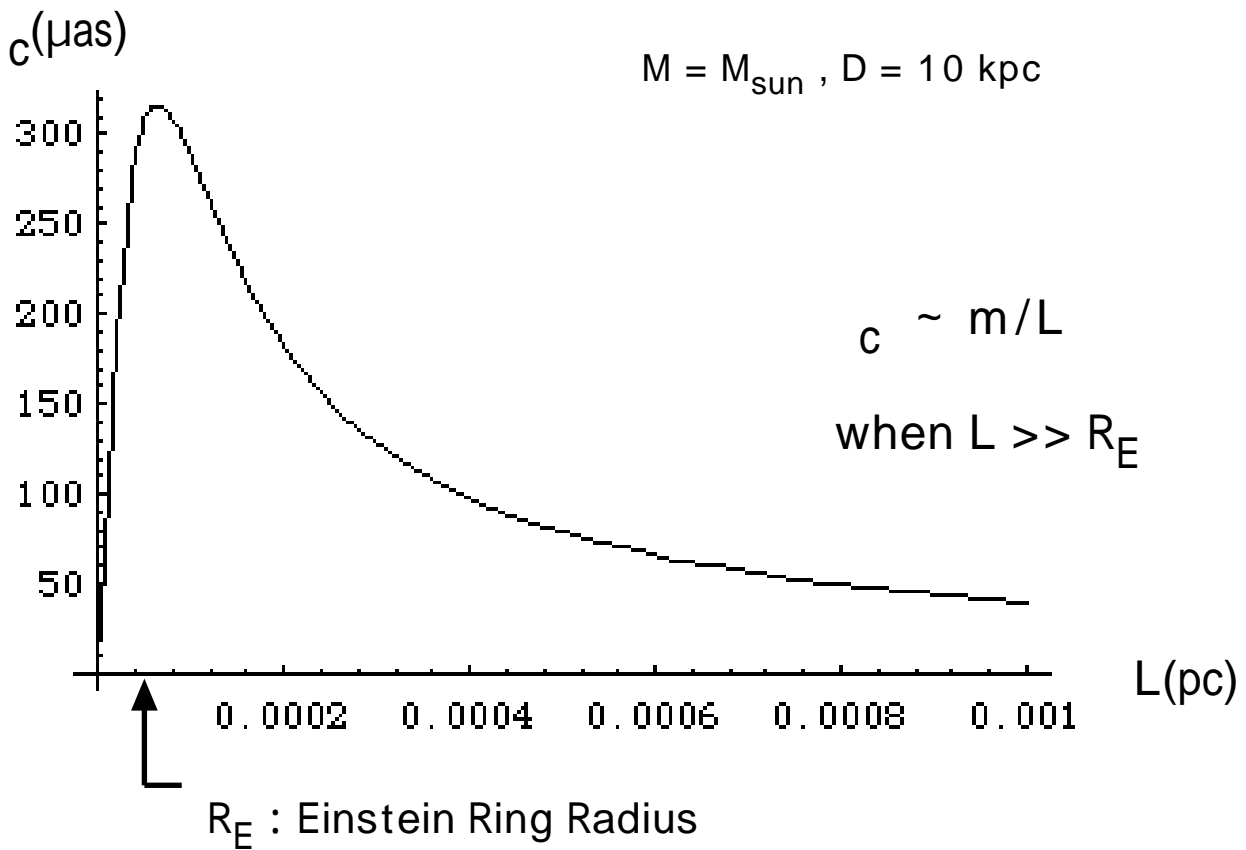
where  $n_D = \frac{\rho_D}{m_D} \delta(m_D)$  ,  $n_M = \frac{\rho_M}{m_M} \delta(m_M)$

## Results of numerical calculation

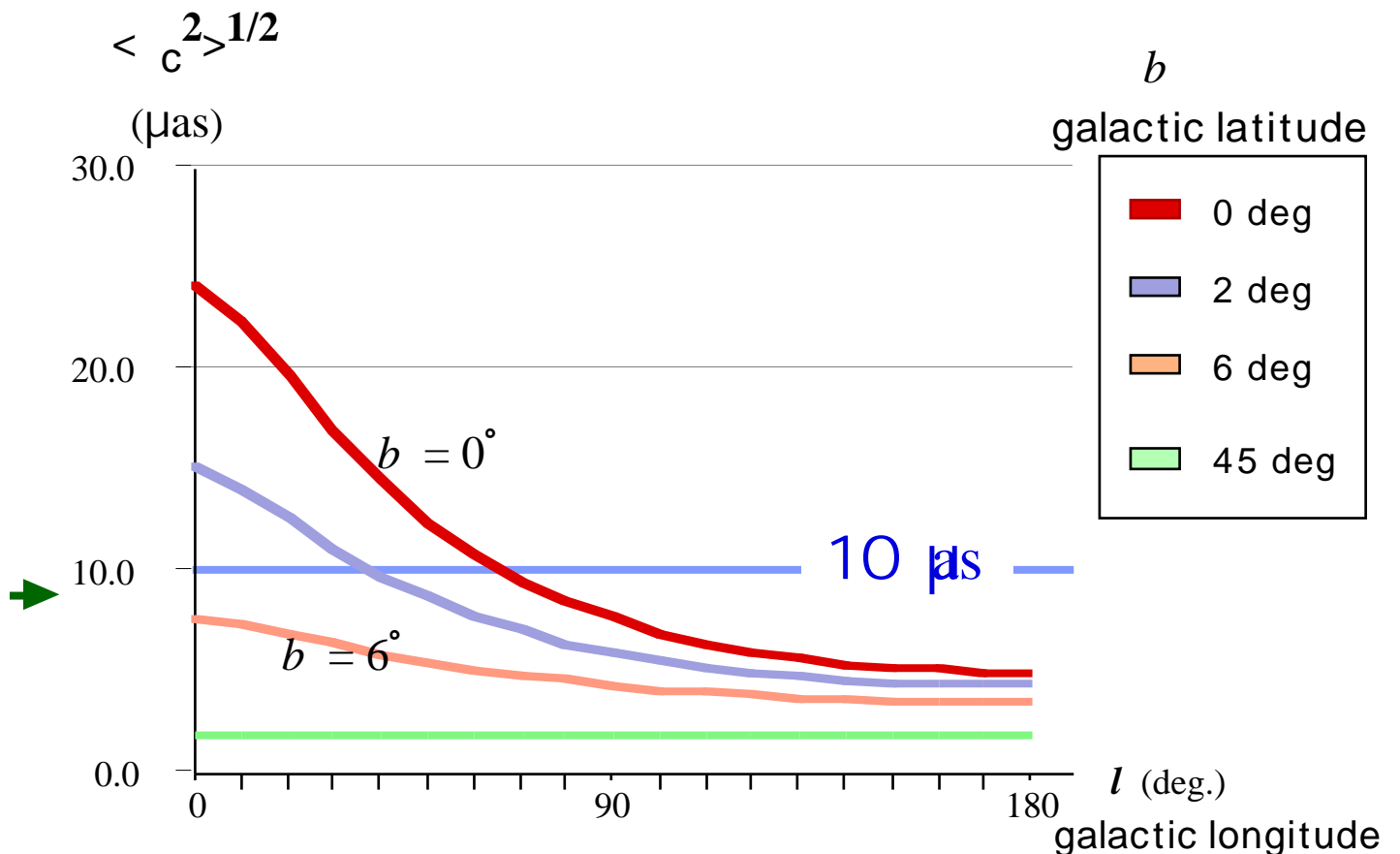
Due to disk stars, mainly.

MACHOs' contribution : very small (under  $0.1 M_{\text{sun}}$  assumption)

## Impact parameter vs. Positional Shift of QSOs



## Expectation Values of Positional Shifts of QSOs



Proper motion of Lensing Matter  
will induce the  
apparent proper motions of QSOs

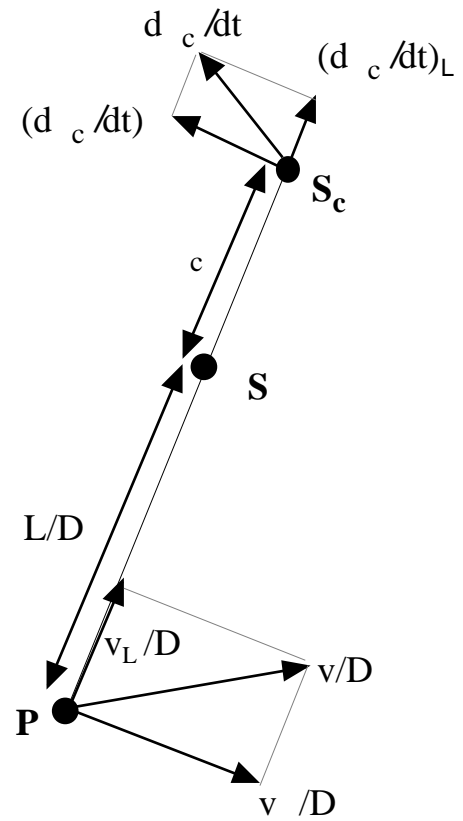
$$\left\langle \left( \frac{d\theta_c}{dt} \right)^2 \right\rangle = \int n(m, \mathbf{r}) \left( \frac{d\theta_c}{dt} \right)^2 dm dV$$

$$= \sum_{i=D,M} \frac{\pi m_i}{3} \int_{D_{\min}}^{\infty} \frac{n_i \langle v_i^2 \rangle}{D} dD$$

Density

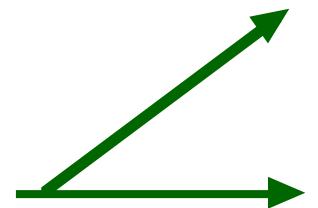
Disk stars : Flat rotation

MACHOs : random transversal velocity



Motion of the Images  
seen from Observer

Results of the numerical calculations



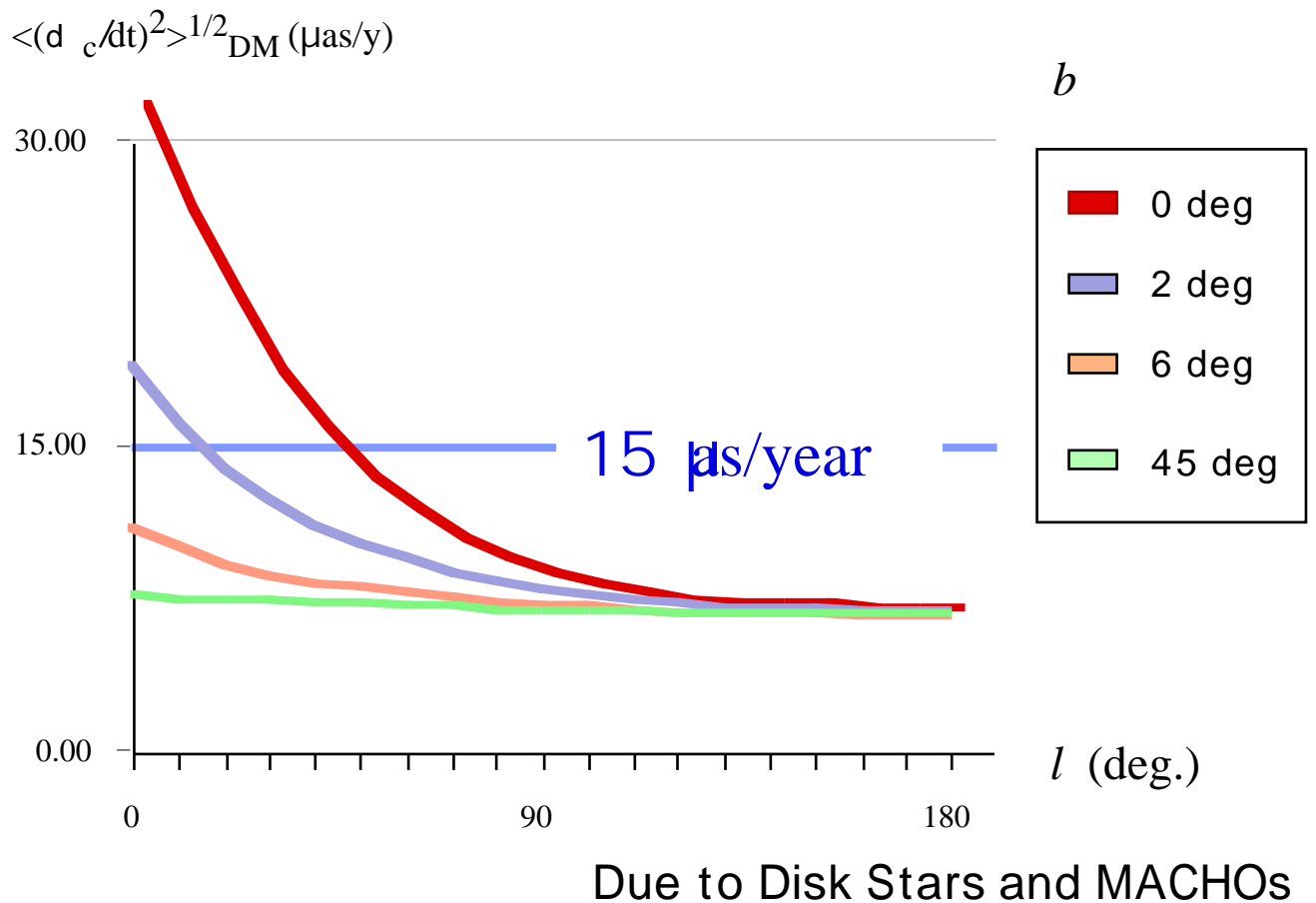
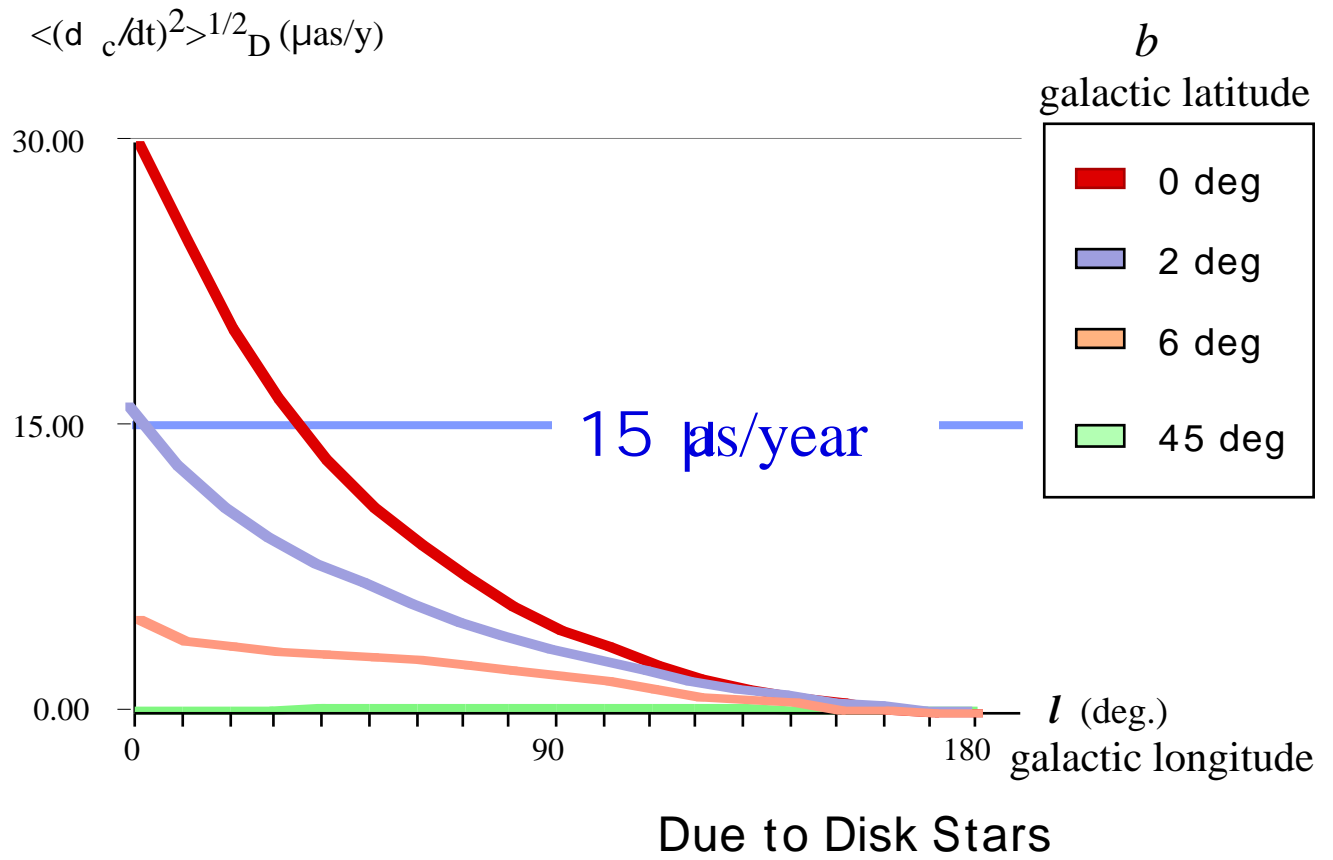
Positional Shifts due to GL      Mass × Density

MACHOs' contribution is small

Induced Proper Motion      Density

MACHOs' contribution is NOT so small ( if exist ! )

# Expectation Values of Induced Proper Motion of QSOs



# Optical depths

Distribution of Pos. Shift and Induced Prop. Motion  
Far from Gaussian

Optical depths are also useful and important

$\sigma_{\text{pos}, i}(\theta_c)$  : Optical depth for positional shift greater than  $\theta_c$

$\sigma_{\text{mot}, i}\left(\frac{d\theta_c}{dt}\right)$  : that for induced proper motion greater than  $d\theta_c/dt$

## Scaling laws

$$\theta_c \sim m/L \text{ and } d\theta_c/dt \sim mv/L^2$$

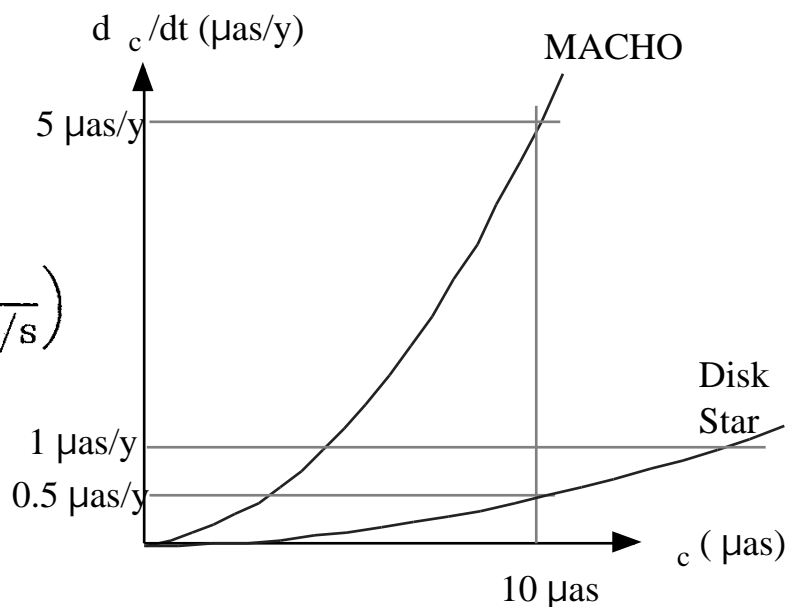
when  $L \gg R_E$

$$\left( \begin{array}{l} \sigma_{\text{pos}, i}(k\theta_c) = \frac{\sigma_{\text{pos}, i}(\theta_c)}{k^2} \\ \text{Accuracy } \times 2 \quad \text{Opt. depth } \times 4 \\ \sigma_{\text{mot}, i}\left(k\frac{d\theta_c}{dt}\right) = \frac{\sigma_{\text{mot}, i}(d\theta_c/dt)}{k} \\ \text{Accuracy } \times 2 \quad \text{Opt. depth } \times 2 \end{array} \right.$$

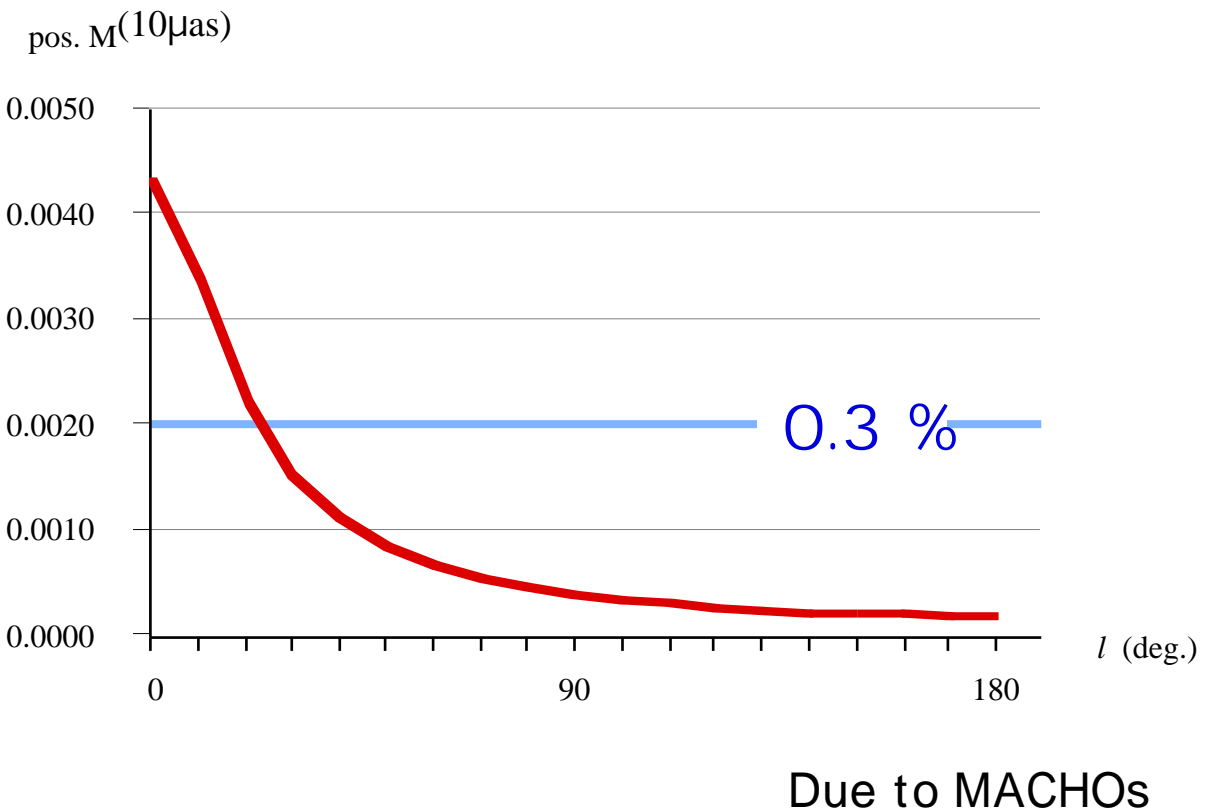
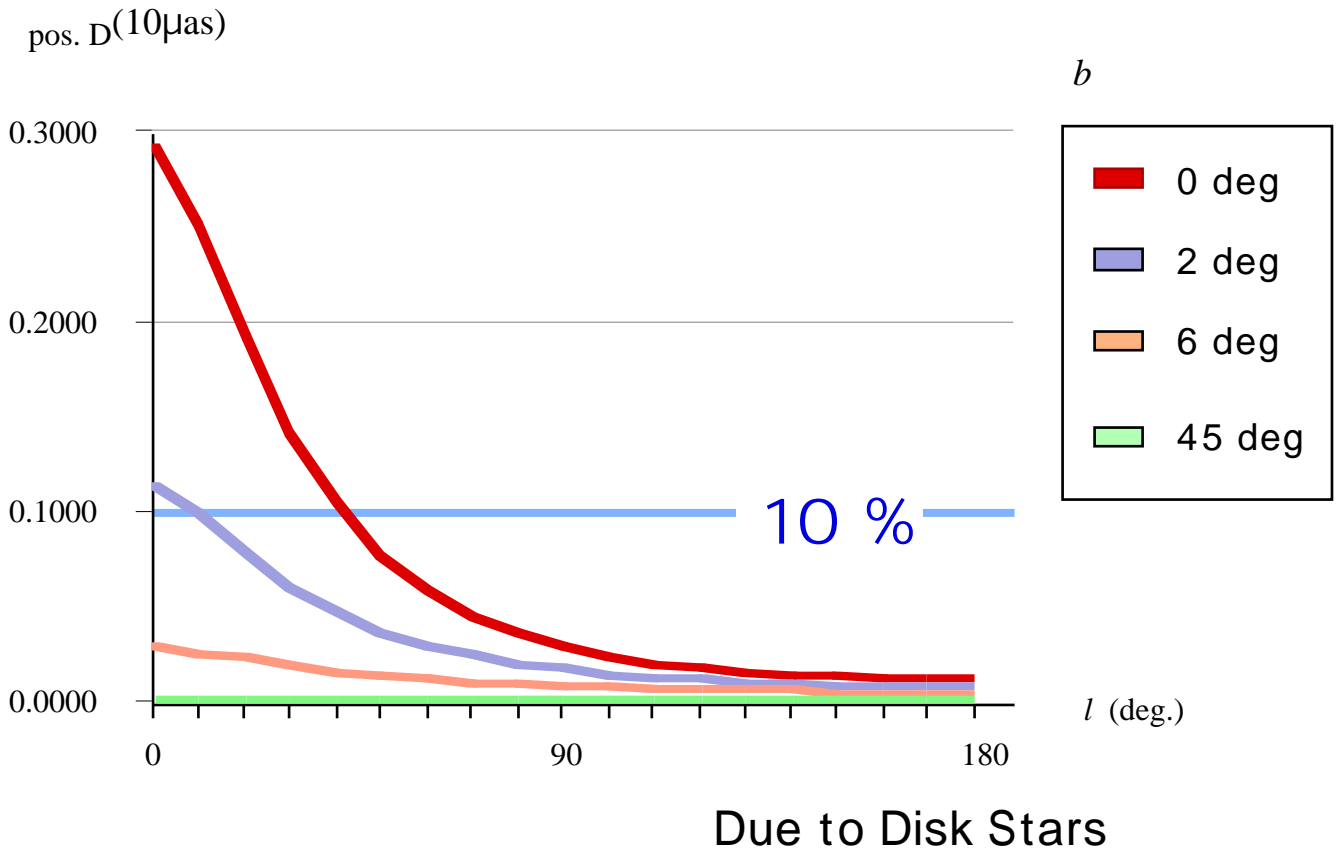
Relation between

$$\theta_c \text{ (}\mu\text{as)} \text{ and } \frac{d\theta_c}{dt} \text{ (}\mu\text{as/y)}$$

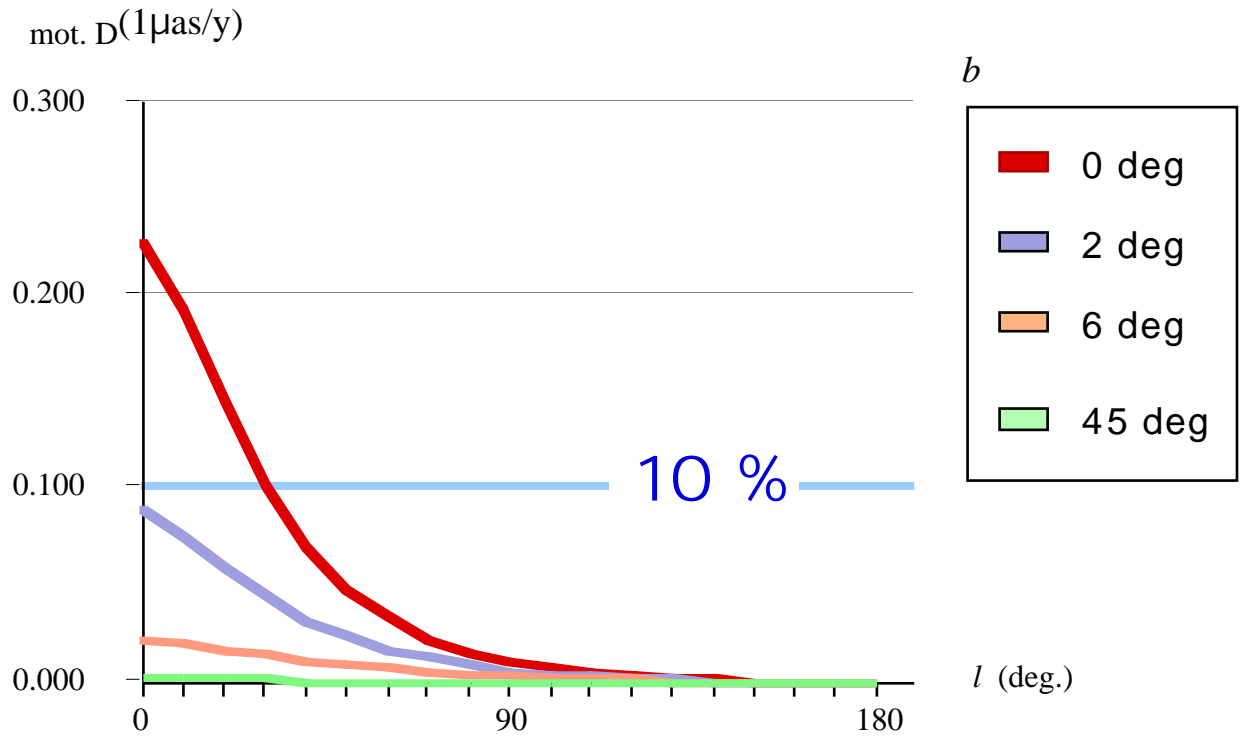
$$\frac{d\theta_c}{dt} = 0.005 \theta_c^2 \left(\frac{1M_{\text{sun}}}{M}\right) \left(\frac{\langle v \rangle}{180\text{km/s}}\right)$$



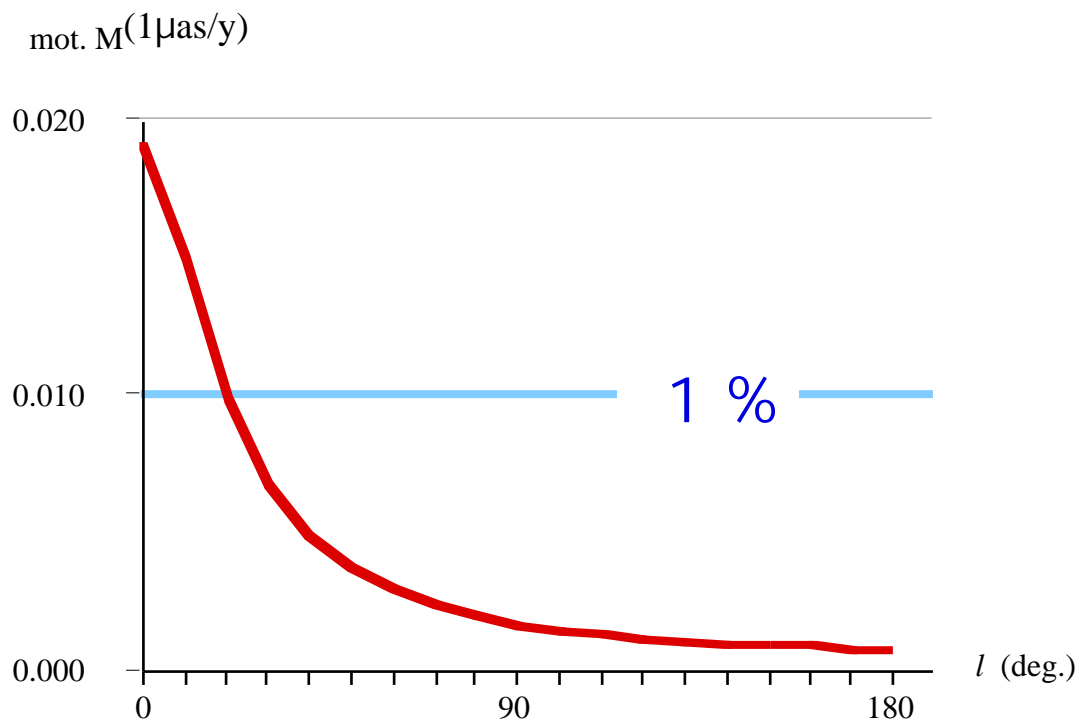
# Optical depth ( $10 \mu\text{as}$ Positional Shift)



# Optical depth ( $1 \mu\text{as/year}$ Induced Proper Motion)



Due to Disk Stars



Due to MACHOs



# Discussions and Summary

## Accuracy limit of Extragalactic Reference Frame

when the optical depths  $\geq 1$ ,  
almost all positions of QSOs are fluctuated,  
affected by multiple sources

Scaling laws tell us that  $\sigma_{\text{pos}}(1 \mu\text{as})$  and  $\sigma_{\text{mot}}(0.1 \mu\text{as/y}) \geq 1$   
in the vast region of the celestial sphere

Difficulty in finding an reference frame  
covering the entire celestial sphere

The AGING of Reference Frame will be observed  
when the QSOs' positions are measured  
with the accuracy of  $10 \mu\text{as}$ ,

New Way to prove the structure of

Disk Stars and Galactic Dark Matter

See the paper (Astron.J. Vol. 114. No.4 October '97)

on the detail of the calculation, MACHOs' core parameter,  
effects of the Orbital Motion of Earth, and

See the poster in S183(Ohnishi, Hosokawa and Fukushima)

MACHOs detection and MACHOs' core parameter