

Intercontinental Comparison of Lattice Clocks using a Broadband VLBI Technique

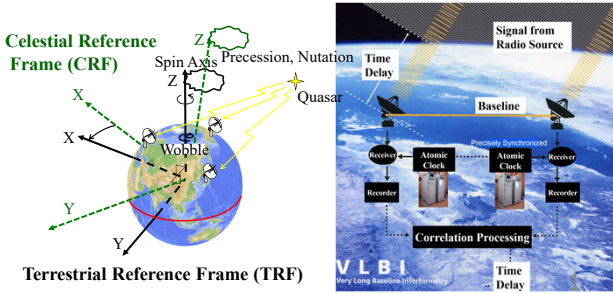
T. Ido, H. Hachisu, N. Nemitz, K. Takefuji, H. Ujihara, E. Kawai, H. Ishijima, M. Tsutsumi, R. Ichikawa, M. Sekido
National Institute of Information and Communications Technology, Koganei, Tokyo, Japan

M. Pizzocaro, F. Bregolin, P. Barbieri, F. Levi, A. Mura, C. Clivati, G. Cerretto, D. Calonico
INRIM, Istituto Nazionale di Ricerca Metrologica Torino, Italy

F. Perini, G. Maccaferri, M. Roma, C. Bortolotti, M. Negusini, R. Ricci
INAF Istituto Nazionale di Astrofisica, Bologna, Italy



What is VLBI? Measurement of spatial coordinates

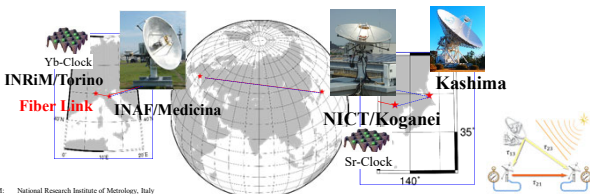


Technique	By means of	Orbit info Dependency	Radio Signal Transmission	Accuracy	Running Cost
TWSTFT	Communication Satellite	Slightly	Need License	< 1.e-15 (Code) < 1.e-16 (Carrier Phase)	>10k USD/yr.
GNSS	GPS /GNSS satellite	Highly	Free	< 1.e-16 (IPPP)	Low
VLBI	Celestial Radio Sources (CRF)	Free	Free	< 1.e-15 (Broadband)

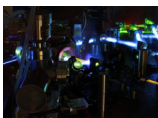
- VLBI is one of space geodetic techniques maintaining spatial reference frames (ICRF, ITRF, and EOP). Time scale can be constructed consistently with these special reference frames.
- VLBI can be operated without using communication satellites. Japan-Europe time link is suffered from lack of communication satellites these years.
- VLBI uses fiducial radio sources in the sky. That leads to long time stability of VLBI measurements.
- VLBI does not need radio transmission license, thus can be used any time anywhere on the earth.
- Higher implementation cost (large antennas) is a disadvantage. This could be overcome by using small diameter transportable stations with Broadband VLBI.

Intercontinental frequency link Experiment: INRIM-INAFA-NICT using VLBI

Target: Intercontinental Frequency Link of Optical Frequency Standard. In addition to existing techniques: TWSTFT, GPS(PPP, IPPP)
 Aug. 2018 :2.4m Antenna exported to INAF/Medicina, Italy



Other components: Two lattice clocks & a fiber link

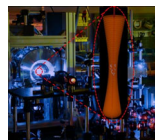


INRIM-Yb1
 (Pizzocaro et al., Metrologia 54, 102 (2017))
¹⁷¹Yb optical lattice clock
 CPM: 518 295 588 500 802 600(36) Hz

Systematic effect:	Correction (ppm)	Uncert. (ppm)
Lattice light shift	-0.84	0.19
2 nd order Zeeman	0.17	0.06
Clock laser light	-0.01	-0.01
SRB shift	23.58	0.14
DC Stark shift	0.2	0.2
Density	0.57	0.06
Background gas	0.05	0.02
Line pulling	0.0	0.1
Sensor error	0.0	0.1
Total	23.72	0.38
Grav. red. shift	-259.97	0.06
SRS		5.00



INRIM-INAFA fiber link
 (Clivati et al., Scientific Reports 7, 40992 (2017))
 Coherent fiber link
 Telescope wavelength 1542 nm
 Length 550 km
 Connects the hydrogen masers at INRIM and INAF with uncertainty < 0.1 ppb/yr



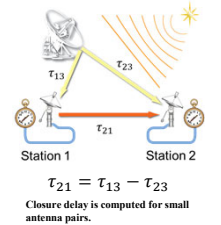
NICT-Sr1
 (Hachisu et al., Opt. Express 25, 9511 (2017))
⁸⁷Sr optical lattice clock
 CPM: 429 228 004 229 773,00(17) Hz

Systematic effect:	Correction (ppm)	Uncert. (ppm)
Lattice light shift	-0.02	0.38
2 nd order Zeeman	0.53	0.03
Clock laser light	0.01	0.01
SRB shift	58.85	0.30
Density	0.01	0.02
DC Stark shift	0.13	0.15
Density	0.13	0.15
Background gas	0.00	0.18
Line pulling	0.00	0.01
Sensor error	0.53	0.15
Total	56.63	0.56
Grav. red. shift	-49.41	0.12
SRS		4.00

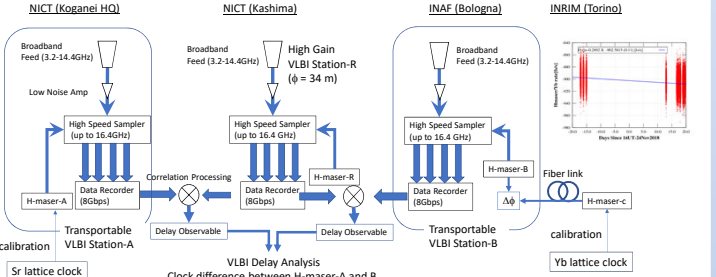
GALA-V Project Overview

Frequency comparison by using transportable broadband telescopes.

- Radio Frequency : 3.2-14 GHz
- Data Acquisition : 4 band (1024MHz width/band)
 - Nominal Freq. Array : Fc=3.7GHz, 5.3GHz, 9.3GHz, 12.1GHz
 - Effective Bandwidth : 3.3GHz (10 times wider than conventional system)



Link scheme

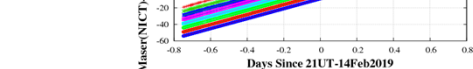


Sr lattice clock
 Clock difference between H-maser-A and B.

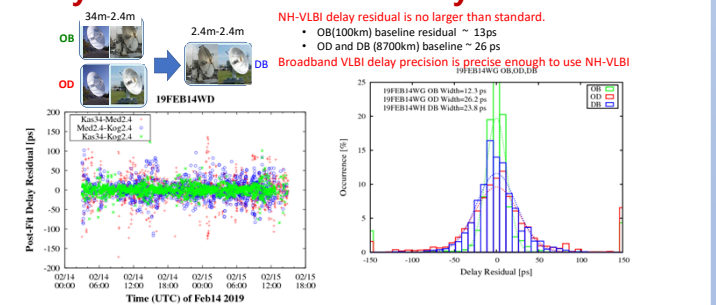
Sr & Yb lattice clock
 Intermittent operations several times per months

VLBI measurement

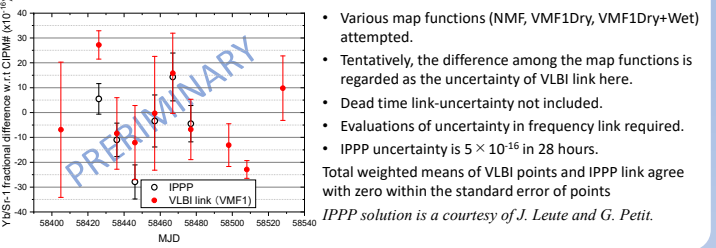
- Correlation analysis indicates the delay between two HMs
- Intermediate Kashima 34m antenna improves S/N
- 36 hours observation, once per 10 days due to the capacity of disk and data transmission
- Bandwidth and polarization synthesis in four bands to improve S/N.



Delay residuals of 'Node-Hub' style VLBI



Result



- Atmospheric delay is the dominant uncertainty in VLBI link
- Various map functions (NMF, VMF1Dry, VMF1Dry+Wet) attempted.
- Tentatively, the difference among the map functions is regarded as the uncertainty of VLBI link here.
- Dead time link-uncertainty not included.
- Evaluations of uncertainty in frequency link required.
- IPPP uncertainty is 5×10^{-16} in 28 hours.
- Total weighted means of VLBI points and IPPP link agree with zero within the standard error of points
- IPPP solution is a courtesy of J. Leuete and G. Petit.

Acknowledgements

IPPP results were computed by **J. Leuete and G. Petit** of **BIPM** using the CNES GINS software.
 High speed research network environment is supported by **JGN,GARR, GEANT, Internet2, and TransPAC**.
 High speed data transfer(~5Gbps) of VLBI data is enabled by **JIVE5ab** developed by **H.Verkoeter**.
 Our project is supported by VLBI Analysis software **Calc/Solve**, Antenna Control **Field System9**, scheduling software **Sked** are developed by **NASA/GSFC**.