



# Astrometry Data Analysis in VERA Project

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# Introduction

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- The final goal of VERA project is to measure the position and proper motion of the Galactic masers with 10 micro-arcsec level accuracy.
- For this purpose, phase-referenced map of each epoch is important.



# Basic Strategy for VERA analysis

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- Finally, we are going to analyze VERA data with our original software.
- Until the development is finished, we use AIPS to analyze VERA data.
- This is mainly because AIPS treats  $(u,v,w)$  data with float-precision (REAL\*4), not double-precision (REAL\*8).

# (u,v,w) Requirement for VERA Analysis

- In order to measure 2 ° separation angles in the error less than 10 micro-arcsec, the requirement of (u,v,w) dynamic range is:

$$\frac{10\mu \text{ as}}{2 \text{ deg}} = 1.4 \times 10^{-9}$$

- The mantissa (仮数部) of float-precision (REAL\*4) usually has 23 bits. Its dynamic range is:

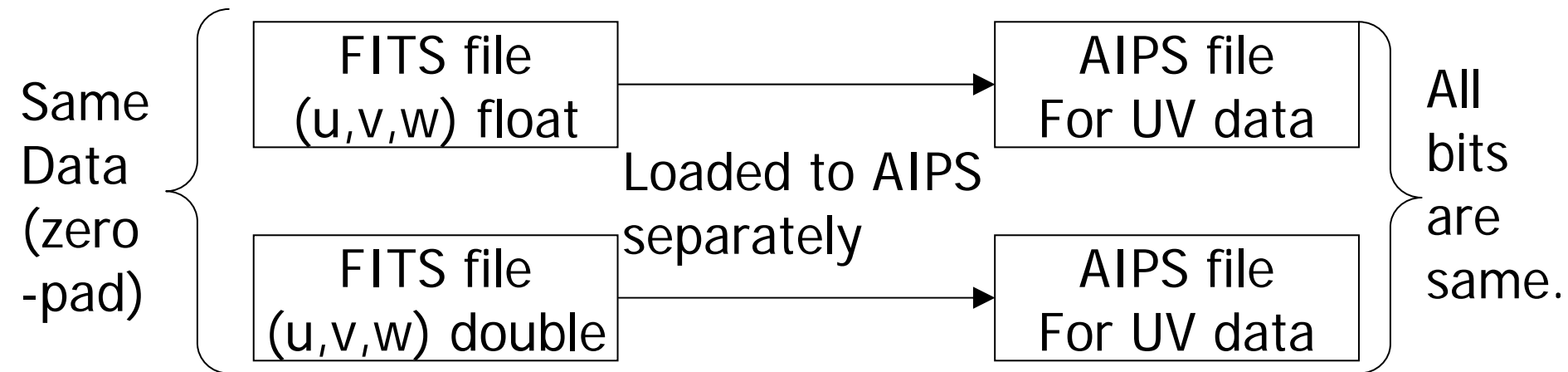
$$\frac{1}{2^{23}} = 1.2 \times 10^{-7} \quad (\text{W. H. Press et al., Numerical Recipes in C})$$

- We must treat (u,v,w) data with double-precision (REAL\*8).

# Importance of Original Software

## AIPS (u,v,w) Test

- Method :



- Result : AIPS accepts double-precision (REAL\*8) (u,v,w) data, but truncates it to float-precision (REAL\*4)



# Importance of Original Software

## – Other Reasons

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- AIPS is not suited to handle huge data automatically.
- AIPS is difficult to use those who are not familiar with it.
  
- By targeting VERA data analysis, we overcome these problems.

# Steps towards VERA

## Astrometry Outputs

- Preprocessing software (ビジビリティ 1 次処理)
    - Calibrations which do not need try-and-error
  - Imaging software
    - From the other calibrations to imaging phase-referenced maps
  - Astrometry software
    - Calculate parallaxes and proper motions from each epoch map
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- We are going to distribute preprocessed data.
  - I hope correlators carry out preprocessing...

# Contents of VERA Astrometry Analysis



- Preprocessing software
  - (u,v,w) re-calculation
  - Path-difference calibration for dual-beam system
- Mapping software
  - Amplitude calibration
  - Bandpass calibration
- Doppler-shift calibration
- Fringe fitting
- Phase-referencing
- Self-calibration
- Imaging
- Flagging
- Time/baseline integration
- Astrometry software






# Basic Strategy for Development

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- Program itself (Command-line interface)
  - C or C++ (Compile with C++)
  - To be developed by VERA team
- GUI
  - Java
  - To be developed by a company (AES).
- Compatible with Mitaka FX correlated database (CODA).
  - In the future, it is compatible with FITS file, too.



# Current Status and Bugs of Original Software

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## ■ Current Status

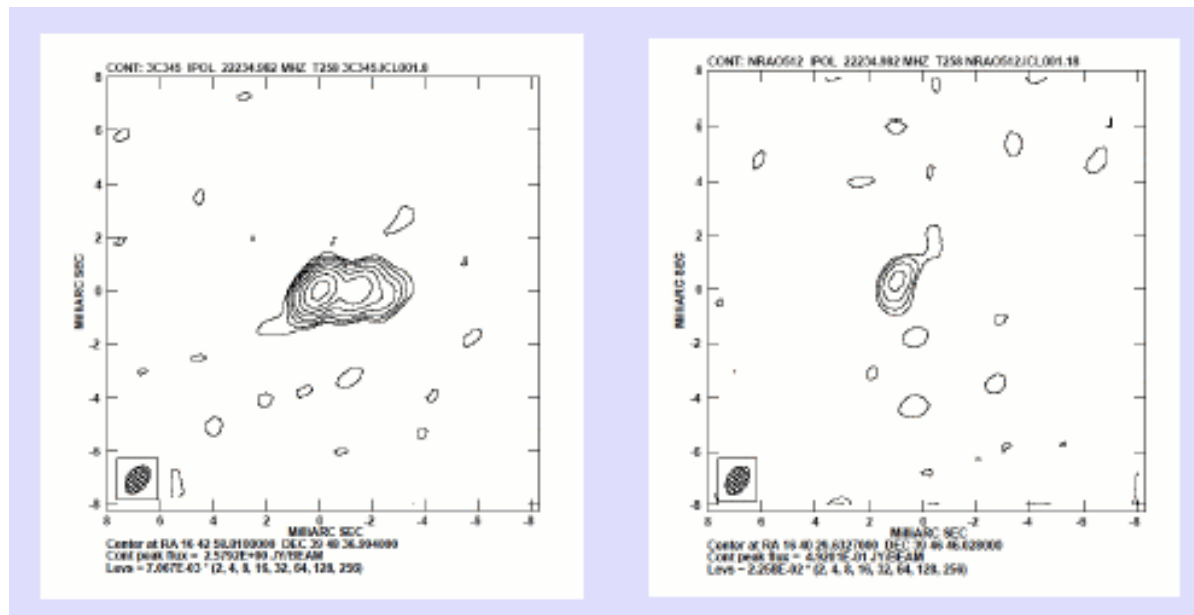
- Preprocessing software : already released
- Imaging software : now developing...
- Astrometry software : now designing...

## ■ Bugs

- Path-difference calibration for dual-beam system :  
Because of some bugs and observation troubles, we do not apply this calibration in preprocessing. We are now applying this calibration at the phase-referencing stage.
- $(u,v,w)$  re-calculation : We are discussing which stage is best to apply it in.

# VERA Data Analysis with AIPS

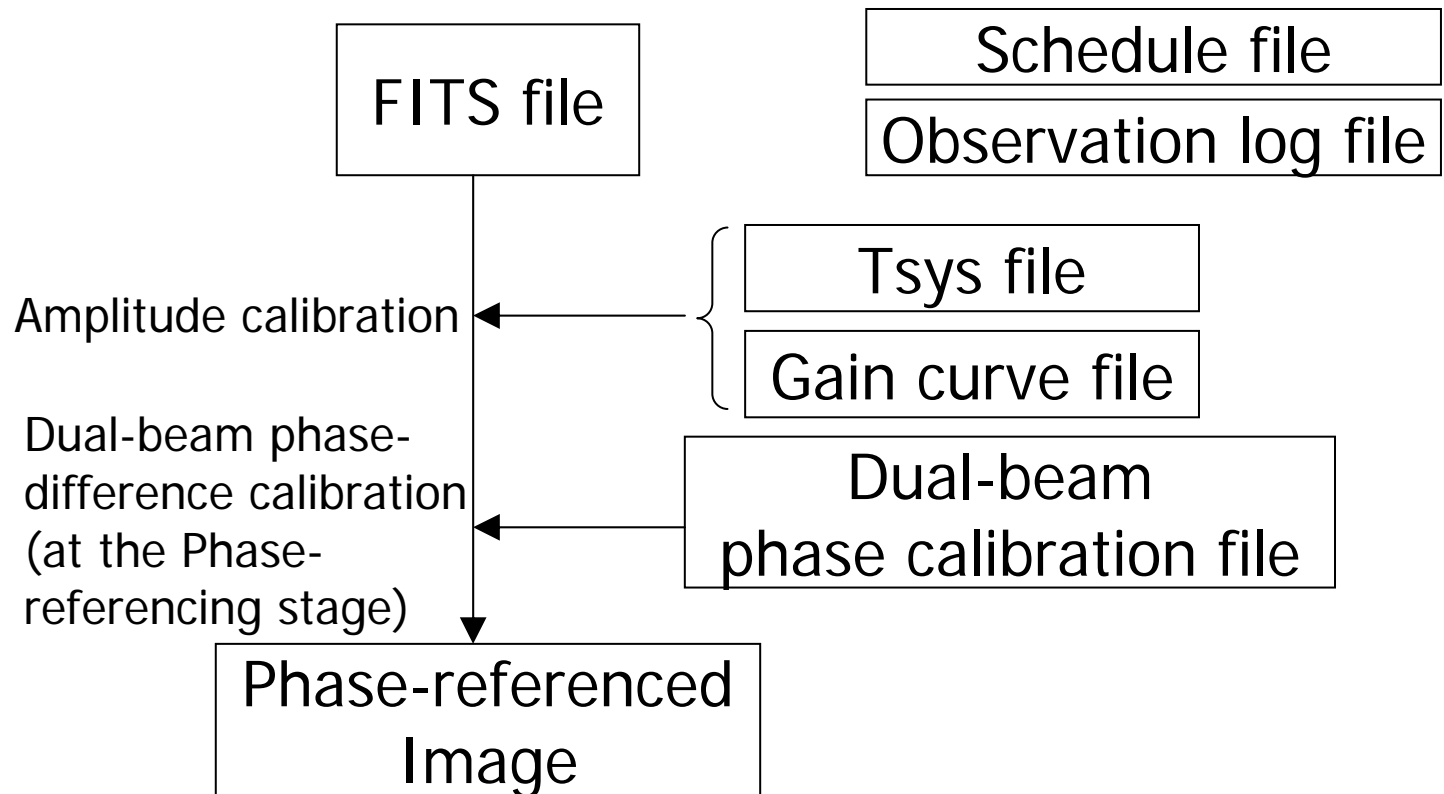
- Mr. Sakakibara obtained phase-referenced map of 3C345 & NRAO512, analyzing with AIPS.



(S. Sakakibara, 2003, Master thesis in Kagoshima Univ.)

# There are Some Extra Files!

- Current status :



- We must manage them.



# Analysis Result Database

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- In order to manage huge numbers of analyses, 'Analysis Result Database' (解析結果DB) is made.
- For the compatibility with AIPS and original software, the scheme of analysis is registered as a text file.
  - For AIPS : RUN file
  - For original software : history file



# Summary

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- Finally, we will analyze VERA data with original software.
  - The reason is the shortage of dynamic range of  $(u,v,w)$ .
  - Basic strategy and current status is reviewed.
- But temporary, we are analyzing VERA data with AIPS.
- We developed Analysis-result database in order to manage the results.