

VERA status and future

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VERA

<http://veraserver.mtk.nao.ac.jp/index.html>

Antenna Diameter 20m (250 μ m)

Observing band 2,8,22,43GHz

Maximum baseline 2273km

Minimum baseline 1000km

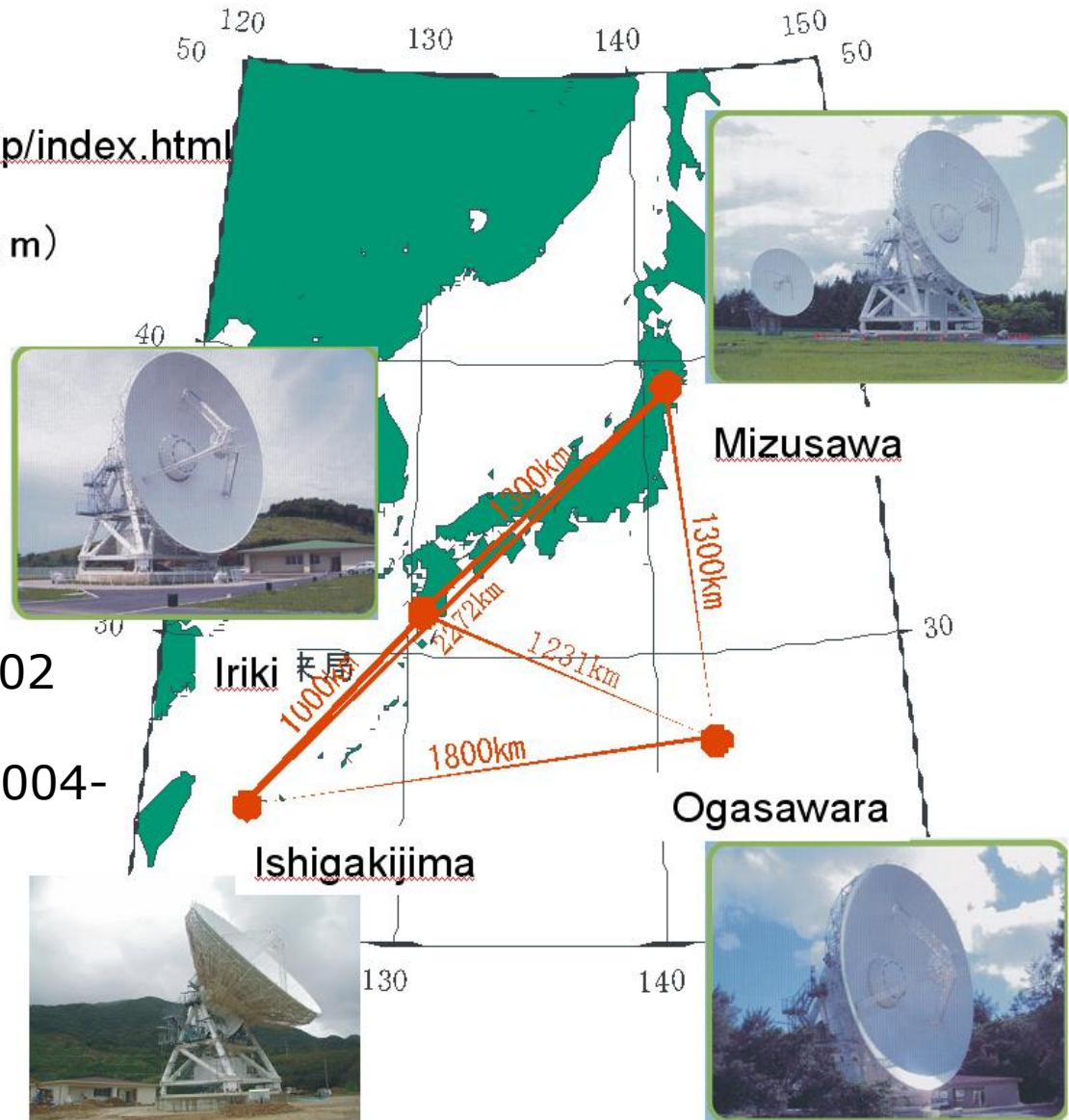
2 beam phase referencing

-> Inst. Path error <0.1 mm

Construction: 2000-2002

Operation: 2002-

Regular observation: 2004-



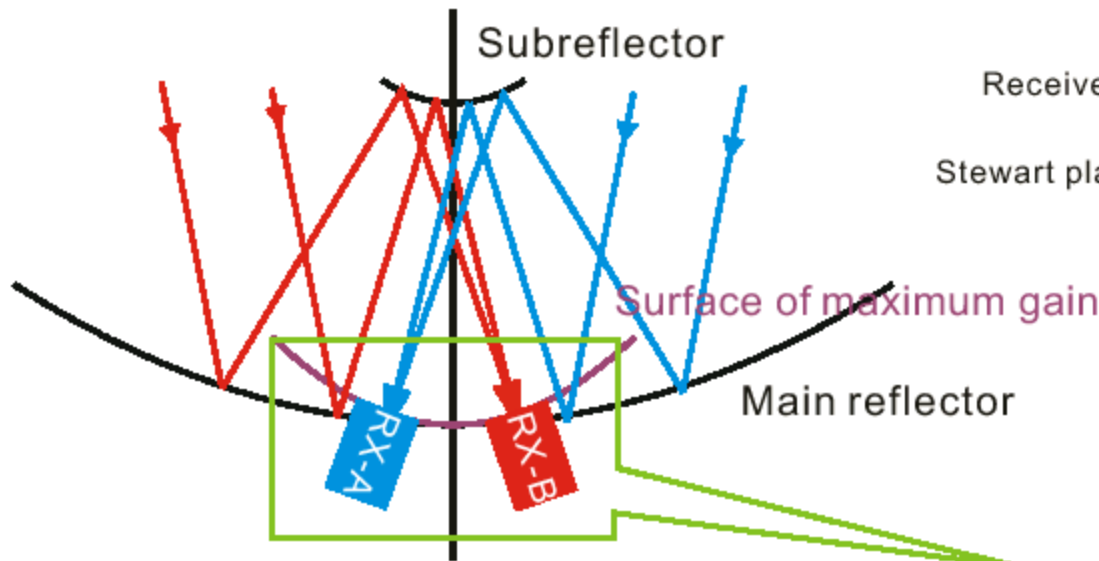
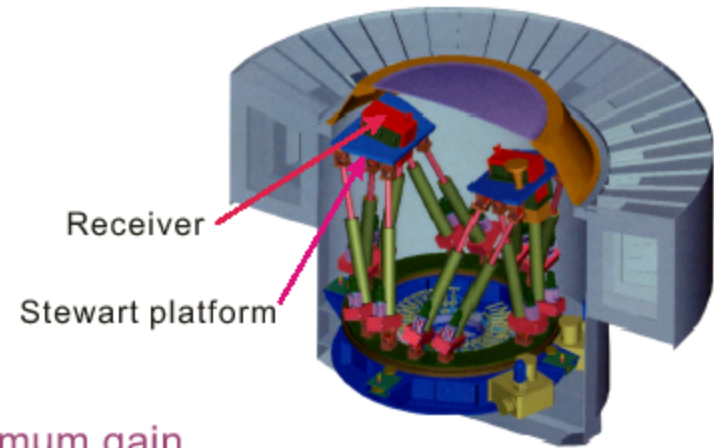
Target source
(maser source)



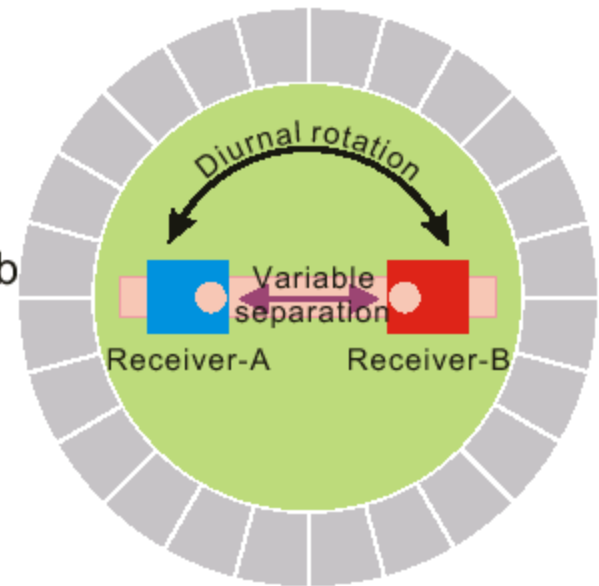
Reference source
(QSO)



Sectional view of the center hub



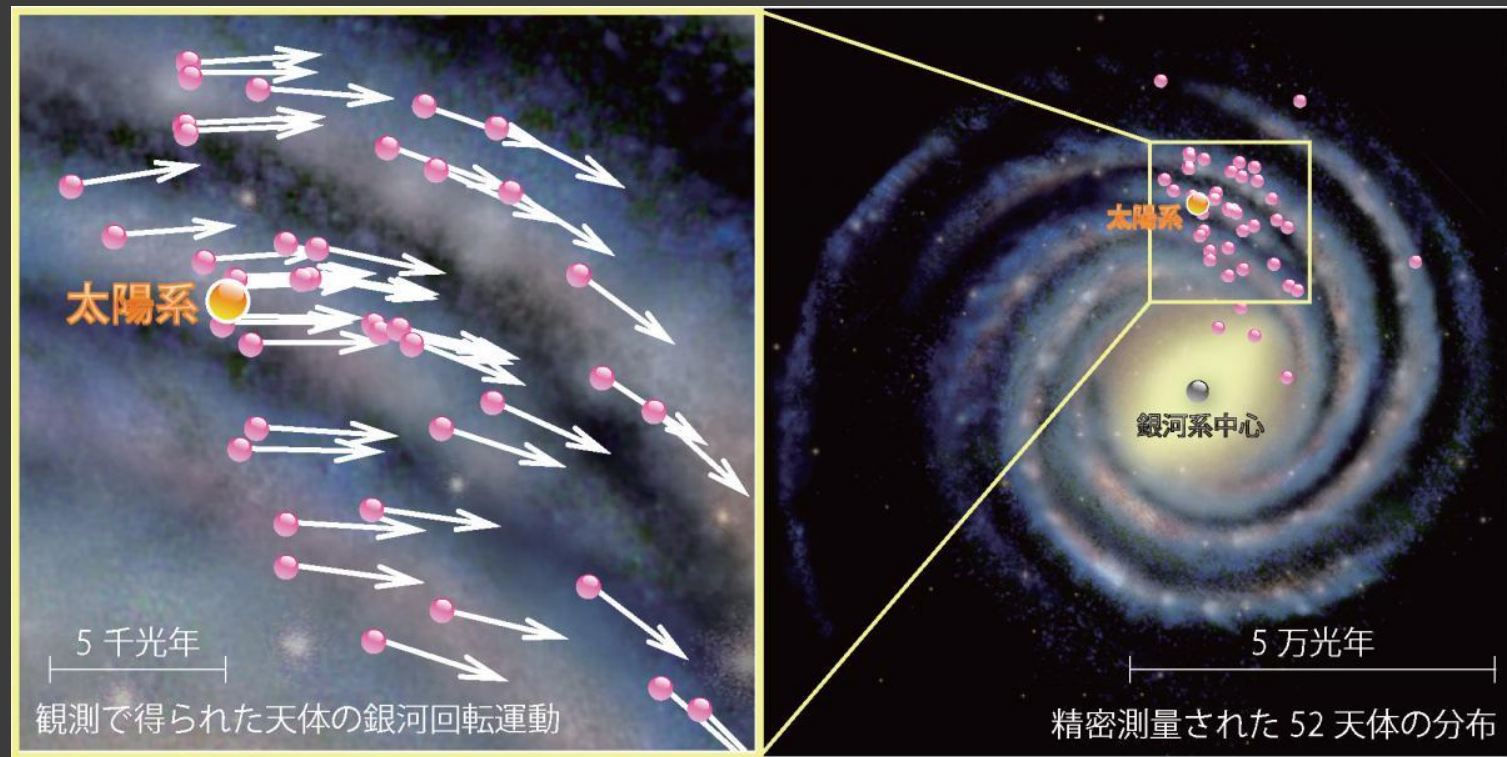
Top view of
the center hub



Concept of the dual-beam
receiving system

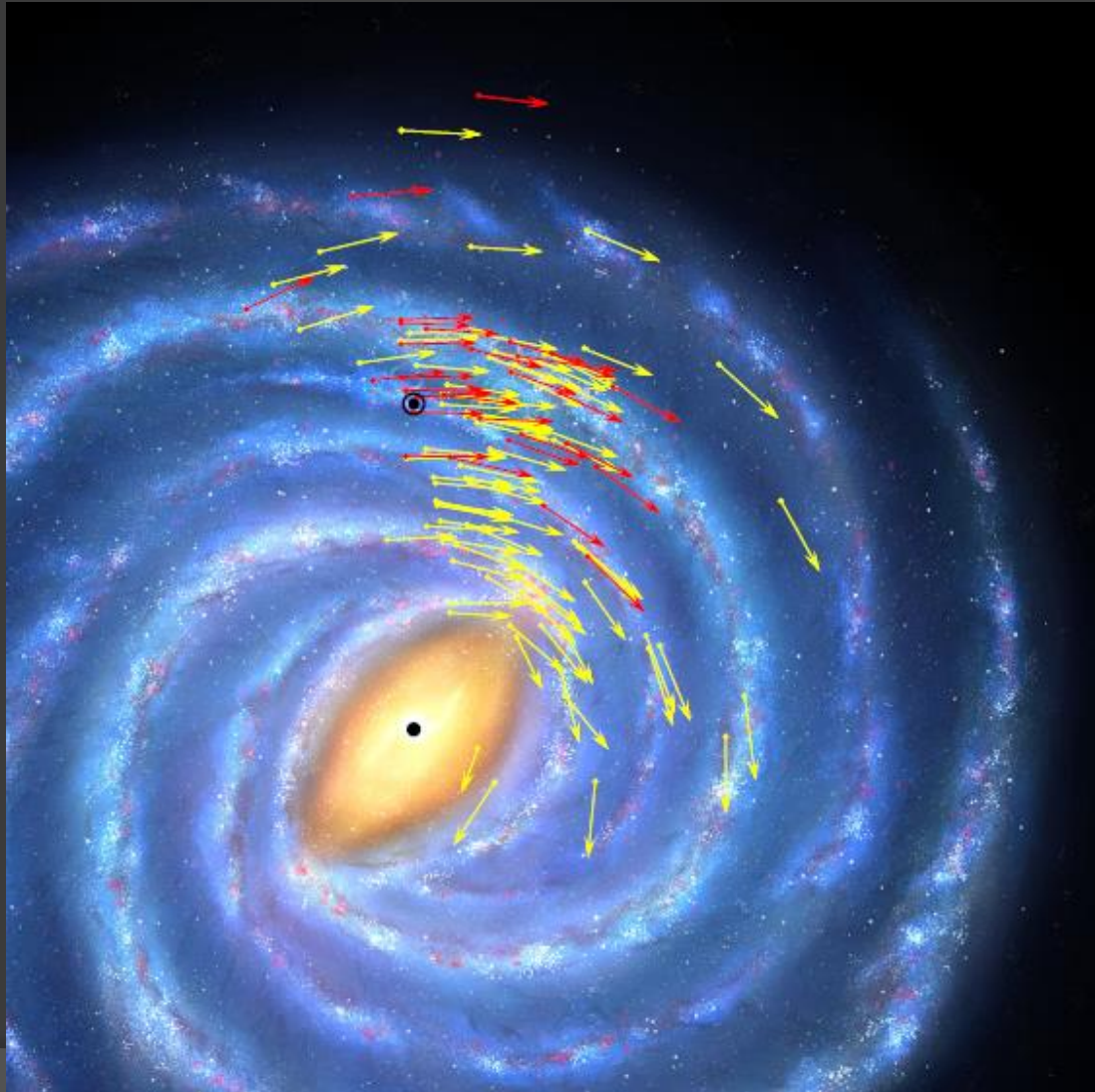
Galactic astrometry in 2012

- Combined with VLBA/EVN data, we have determined Galactic parameters using 52 sources.



Galactic parameter can be determined by using more than 50 sources with VLBI astrometry

VLBI astrometry results for Galactic masers



→ VERA
→ VLBA&EVN

Revising Galactic constants

New (preliminary)
(Honma+12)

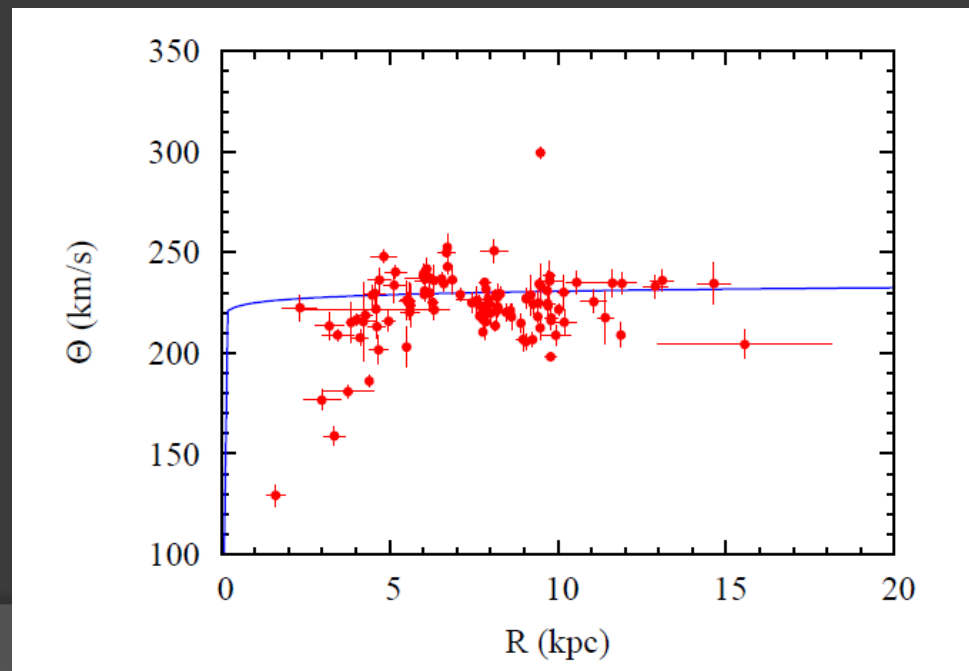
Previous

N_{src} 114 sources

(52 sources)

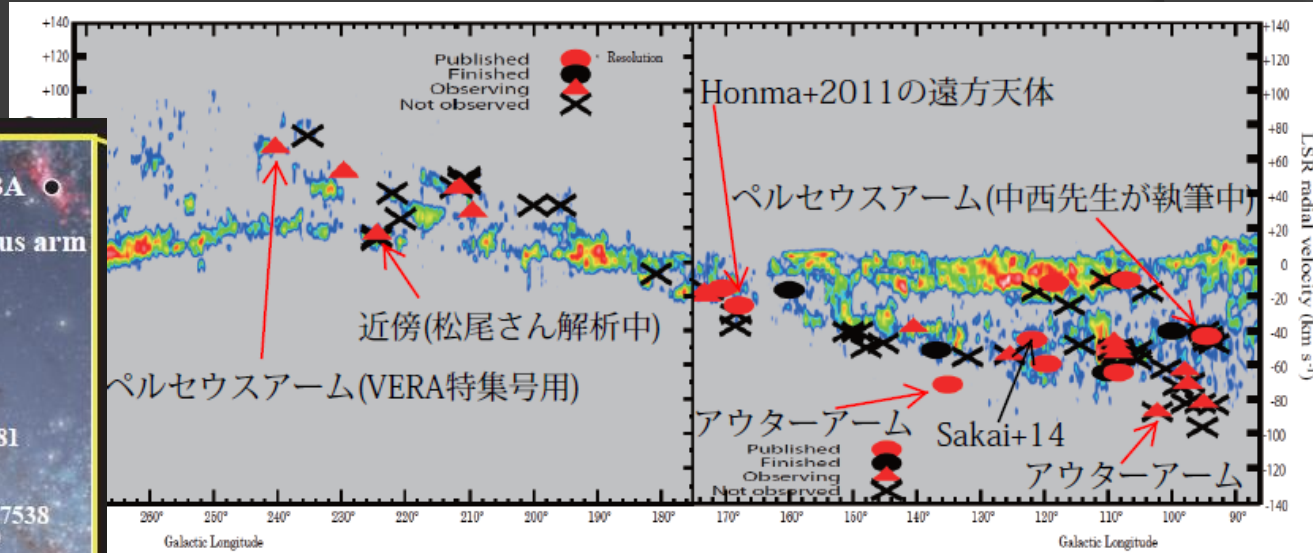
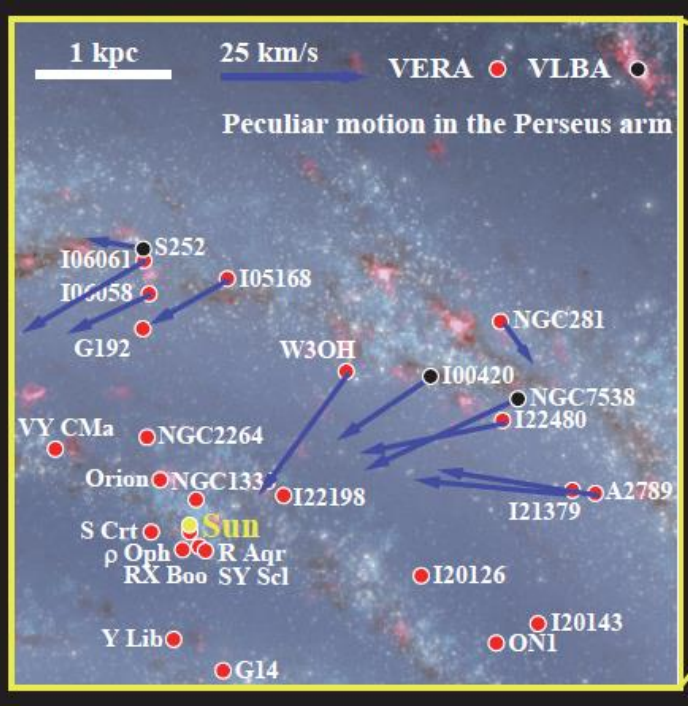
$R_0 = 7.95 \pm 0.21$ kpc (8.05 \pm 0.45 kpc)

$\Omega_0 = 28.95 \pm 0.43$ kpc
(29.57 \pm 0.78
kpc)



Spiral Arms

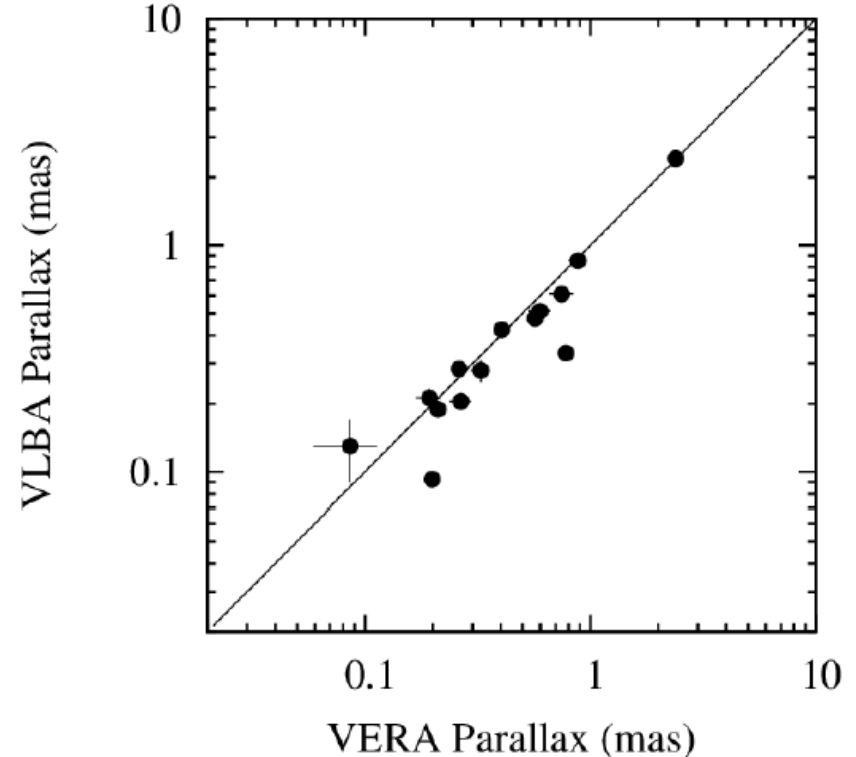
Location and non-circular motion in spiral arm



Outer galaxy objects (in 1-b diagram)

VERA vs VLBA in parallax

- ~10 sources observed with both VERA and VLBA (in some cases different bands)
- Generally consistency, no systematic offset
- Discrepancy seen in a few sources. Needs to be checked.

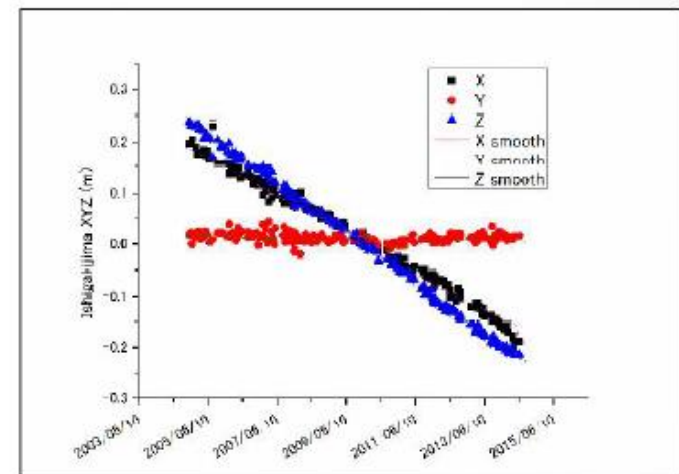
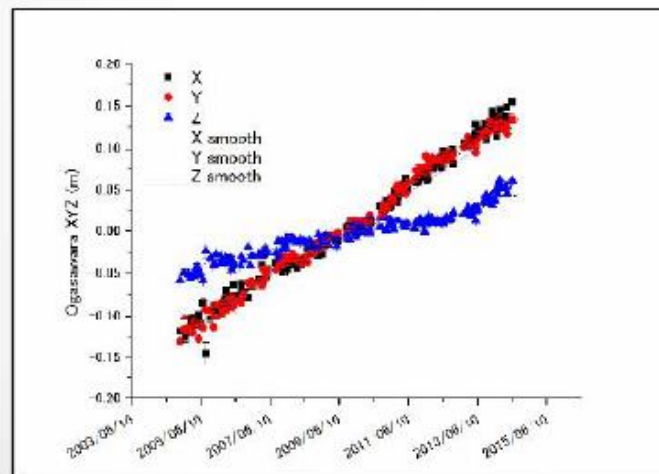
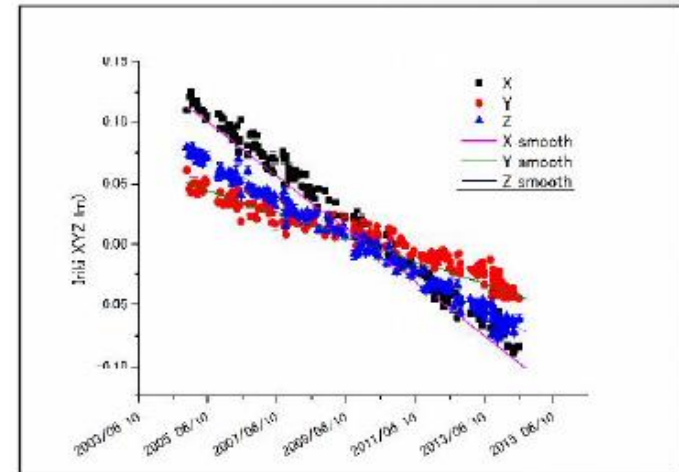
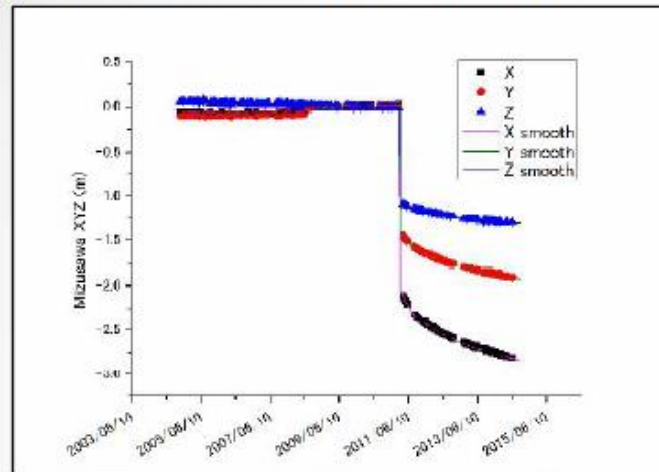


VERAとVLBAの視差の相関図

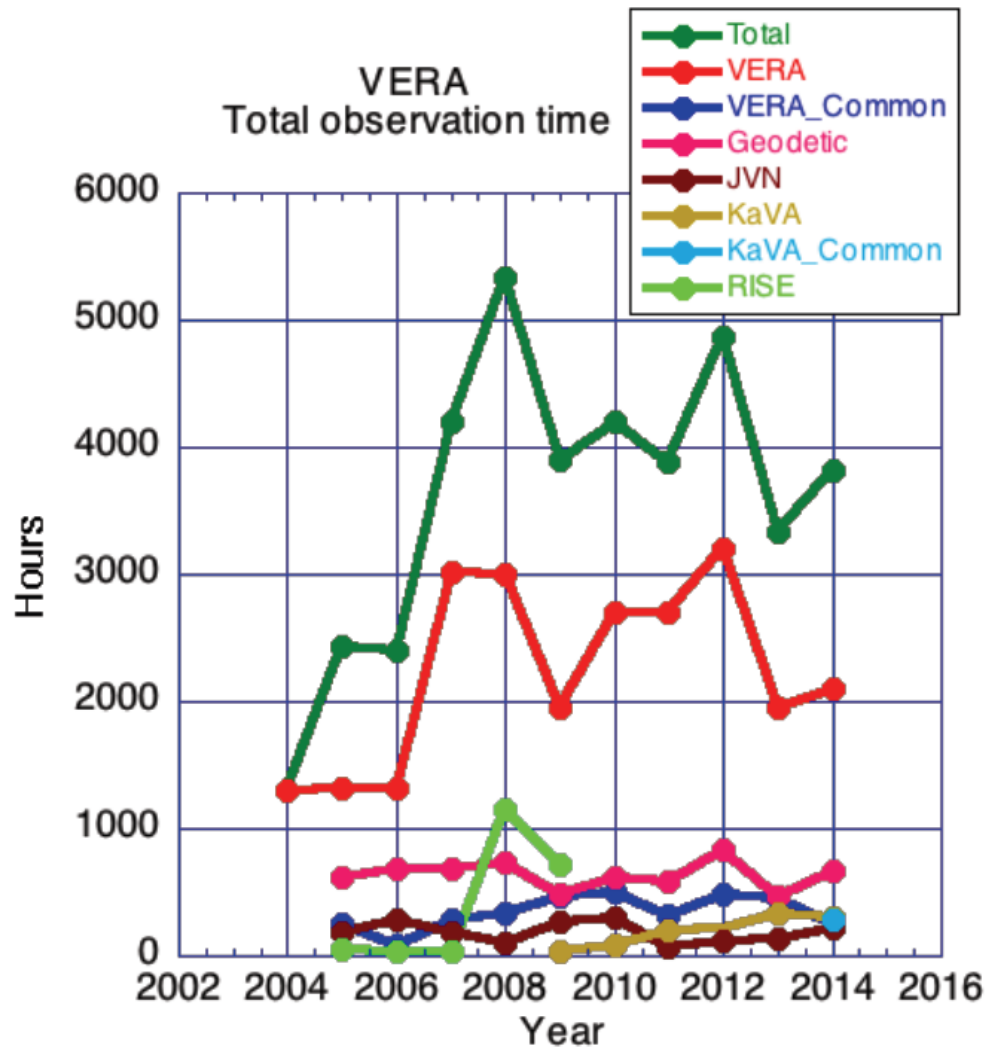
Motion of VERA station

地心座標系でのVERA各局の座標変化

Epoch=2010.0の座標に対する差



Operation time of VERA

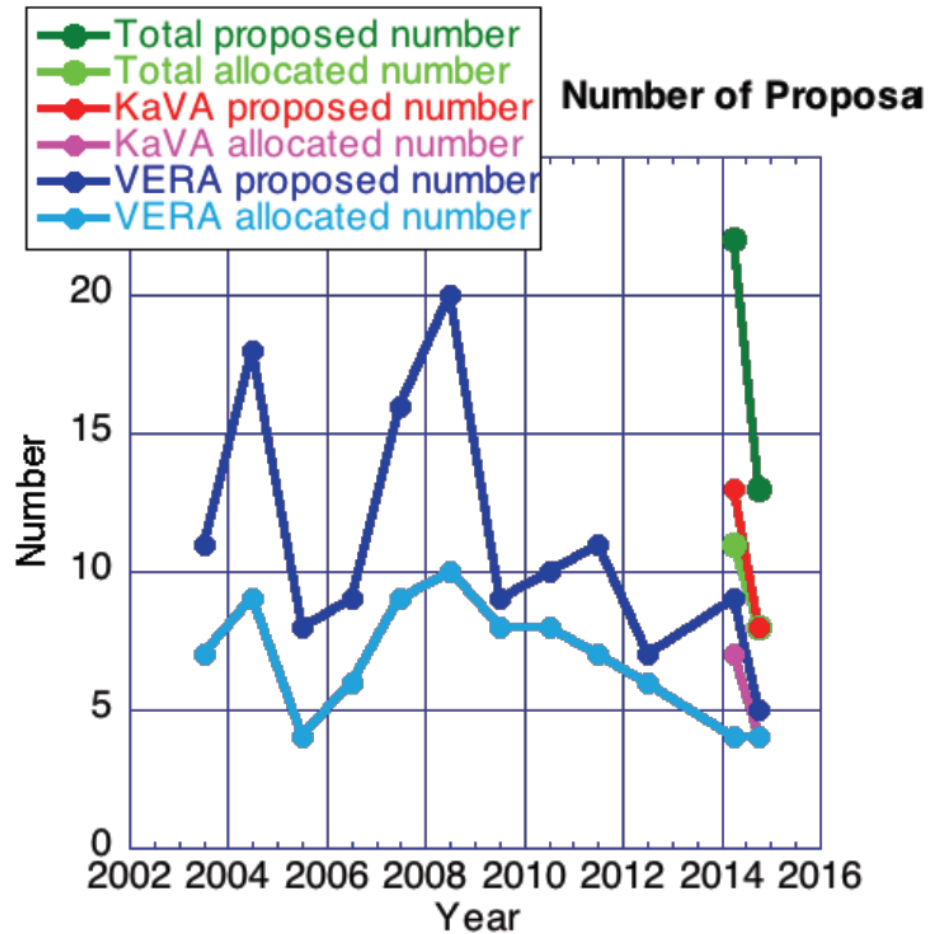


Open use

- ◎ VERA with Nobeyama(45m) and Kashima(34m) is opened for global users.
 - Open time is 700 hours/year including KaVA.
- ◎ KaVA (KVN and VERA array) is opened to East-asian astronomy communities at current stage, which will be open to the world near future.
 - Current open time is 500 hours/year.
- ◎ Deadline of proposal is twice per year.

<http://veraserver.mtk.nao.ac.jp/restricted/index.html>

Number of proposals for VERA and KaVA



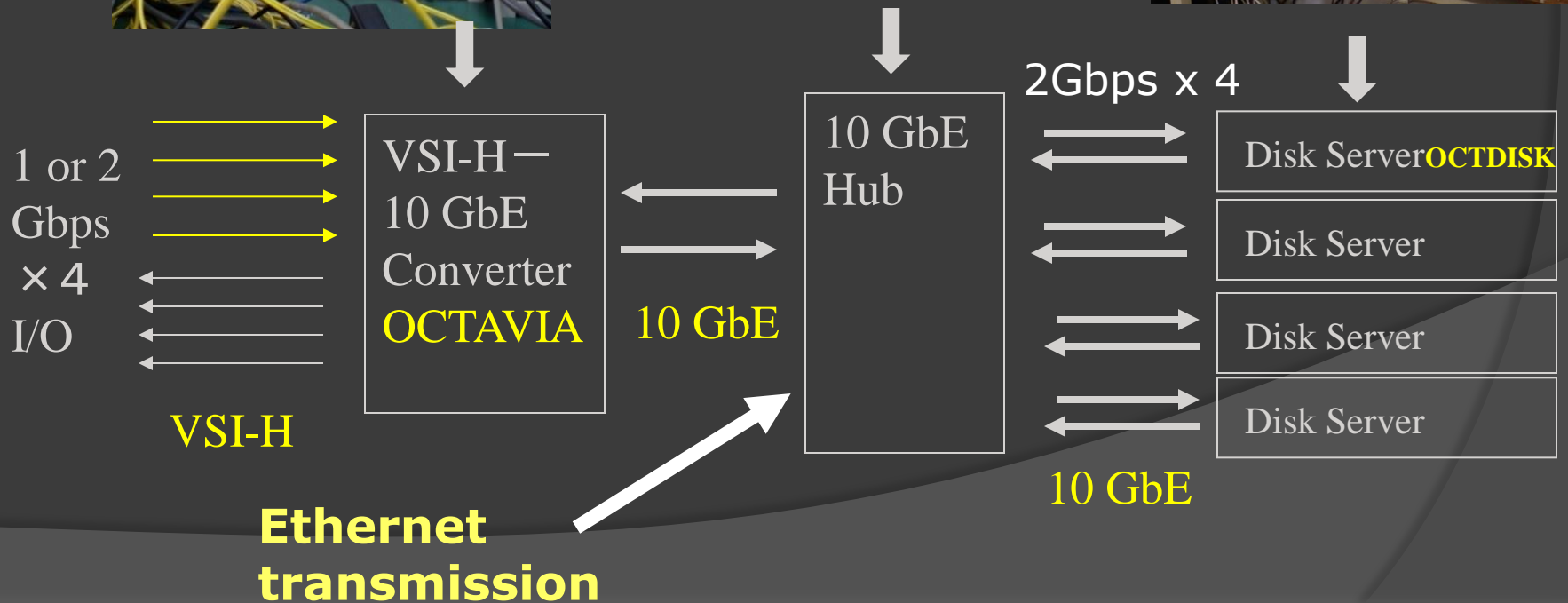
Developments

- ◎ **High speed data recording using disk system**
 - Usual operation of OCTAVE system + Soft Corr (1~8Gbps) will be started at 2015.
 - VSREC (12Gbps by PC recording) is under developing.
- ◎ K/Q dual polarization is under testing.
- ◎ Ultra high speed sampler (<50GHz) is under studying.
- ◎ K/Q simultaneous receivers are under studying under the collaboration with KVN.

OCTAVE

(8Gbps Disc recorder)

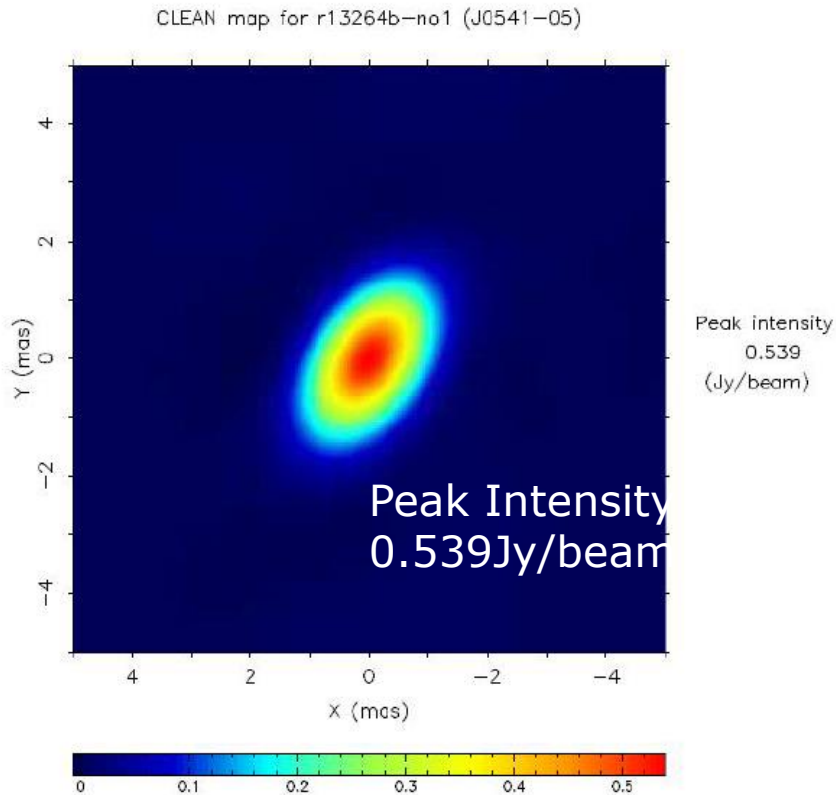
-> 3 times sensitive



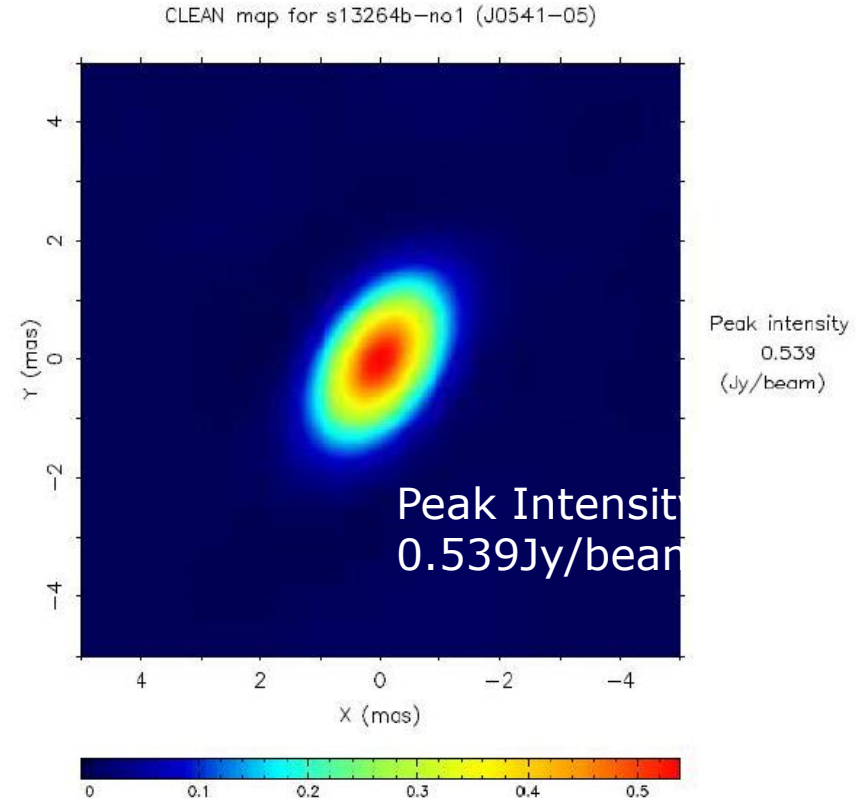
New observing mode by using Disk systems

	Recording rate	Recorder	HDD	Correlator	KJJVC	2beam (phase-ref)
VERA (KaVA)	1 Gbps	OCTADIS K	All	Mitaka-FX Soft corr	◎	○
VERA	4 Gbps (A=B=2G)	OCTADIS K	80h	Soft corr	○	Under testing
VERA (test)	12 Gbps (A=2G, B=10G)	VSREC	100h	Soft corr	○	Under development
NRO45	4 Gbps (1Gbps)	OCTADIS K	50h	Mitaka-FX Soft corr	○	
JVN-OCTAVE	2 or 4 Gbps	OCTADIS K K5VSI		Soft corr	○	

Verification of software correlator



Tape > Mitaka-FX

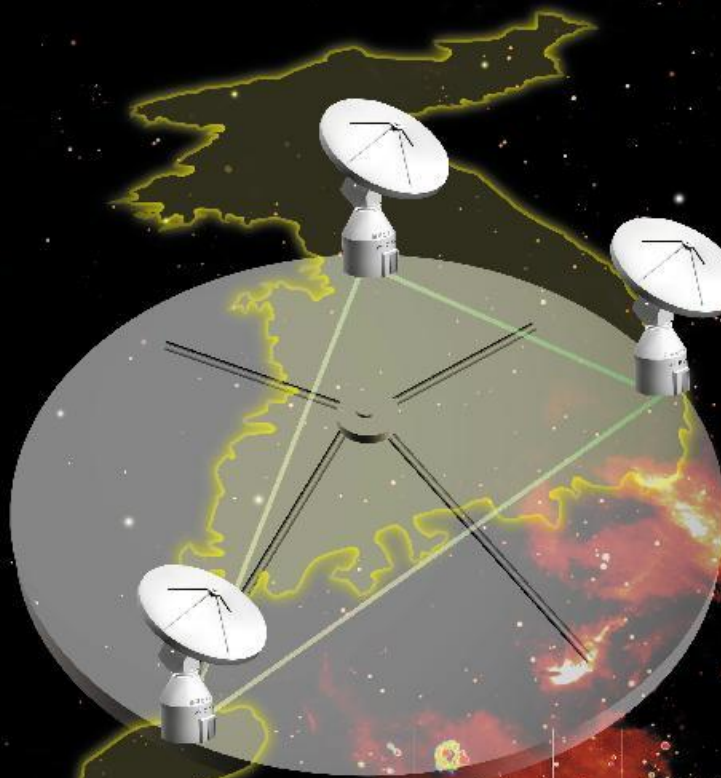


Tape > Soft corr

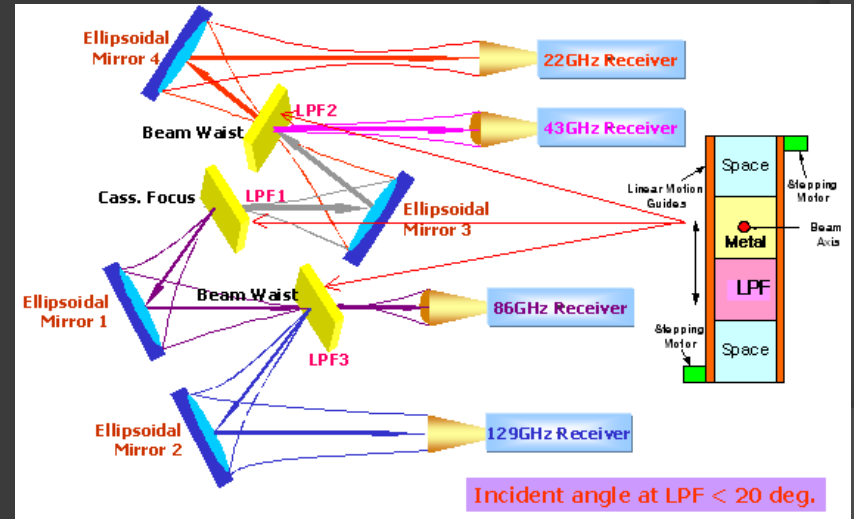
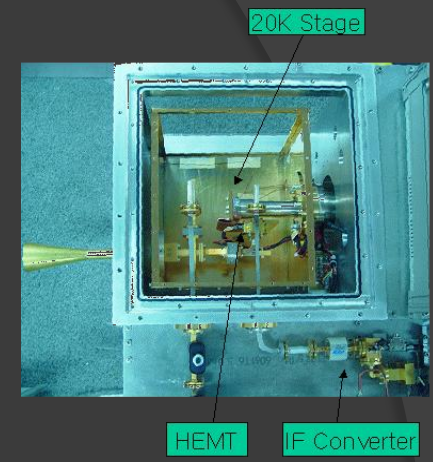
Correlator

- Software correlator is almost verified with comparison current hardware correlator(Mitaka-FX).
- It will be moved to Mizusawa until 2015 Mar. and started regular operation.
- Mitaka-FX correlator will be closed in 2015 which means no correlation of tape based recording such as DIR1000 and 2000.

KVN 한국우주전파관측망 Korean VLBI Network



한국천문연구원
내·외신파안문대 KVN사업부



Constructions of three stations were completed on Dec. 2008 !

Simultaneous Multi-Frequency Obs. -
Phase Compensation, mm-VLBI

East Asian VLBI Network array



Future Plan

- ◎ NAOJ Director General requires the intensive review of VERA project at 2022, which should consider the termination or continuation of VERA program.
- ◎ After 2023, Mizusawa VLBI observatory is required a new project including VERA continuation.
- ◎ Possibility of future plans are
 - East Asian VLBI network will extend to Pacific and Global array.
 - Joining SKA

Summary

- ④ VERA is powerful tool for astrometry and made a significant contribution for Galactic parameter determinations.
- ④ High speed disk recording and correlation system will be operational shortly and more wide band system is under developments.
- ④ KaVA observation is operational and EAVN is testing.
- ④ Discussion of future plan is just started seriously. It is also needed with the Japanese VLBI community.