

The IVS and its VLBI Global Observing System VGOS

Axel Nothnagel

Institute of Geodesy and Geoinformation, University of Bonn

and

many colleagues of the IVS

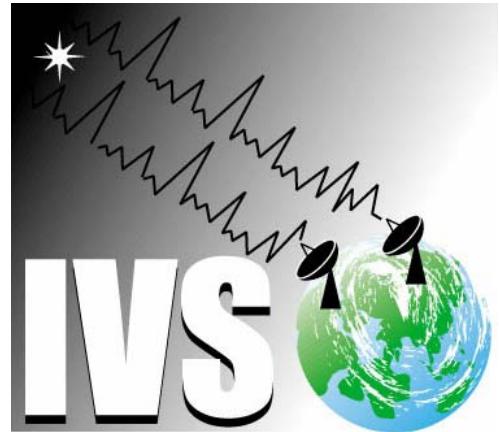
International VLBI Service for Geodesy and Astrometry

Contribution to IAG's Global Geodetic Observing System (GGOS)

- 1 mm/0.1 mm/y
- continuous
- stable over decades
- global distribution



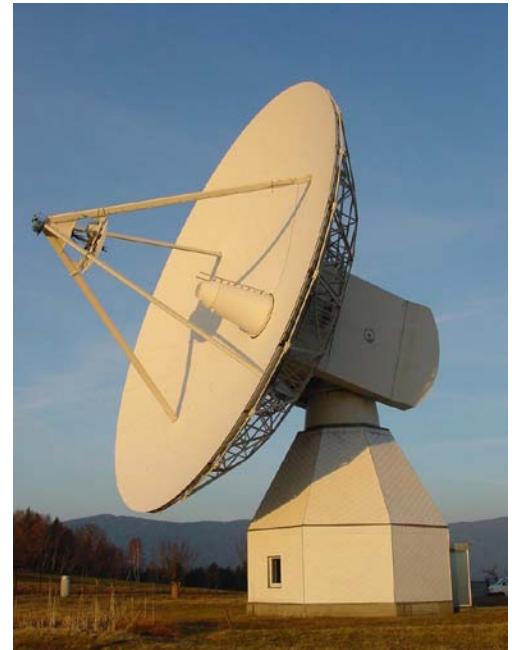
International
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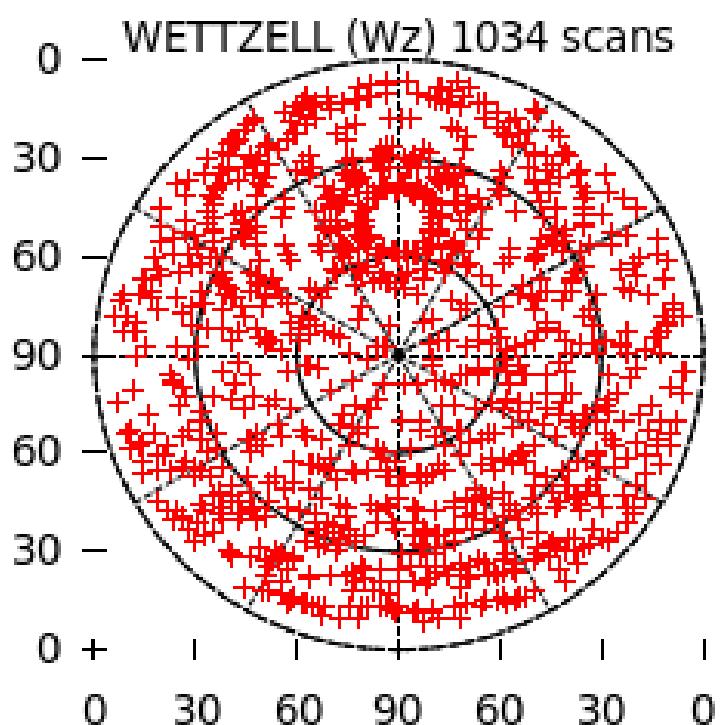
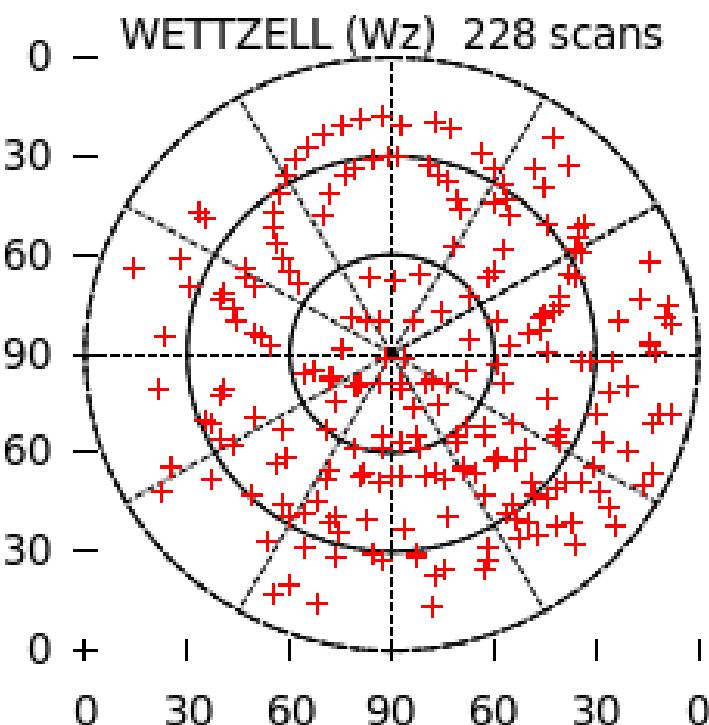
- Developments towards VGOS
VLBI(2010) Global Observing System
 - Telescopes, Correlation, Analysis
 - Organisational aspects



Current IVS Station Network



- New generation VLBI infrastructure
 - dense sampling of local sky for optimal estimation of atmosphere parameters



- New generation VLBI infrastructure
 - dense sampling of atmosphere
 - agile telescopes
→ small (12 – 13 m)
12°/sec
 - up to 2 observations per minute
(2880/day)

$$\sigma_\tau \propto \sqrt{\frac{1}{A_1 A_2 \cdot B}}$$

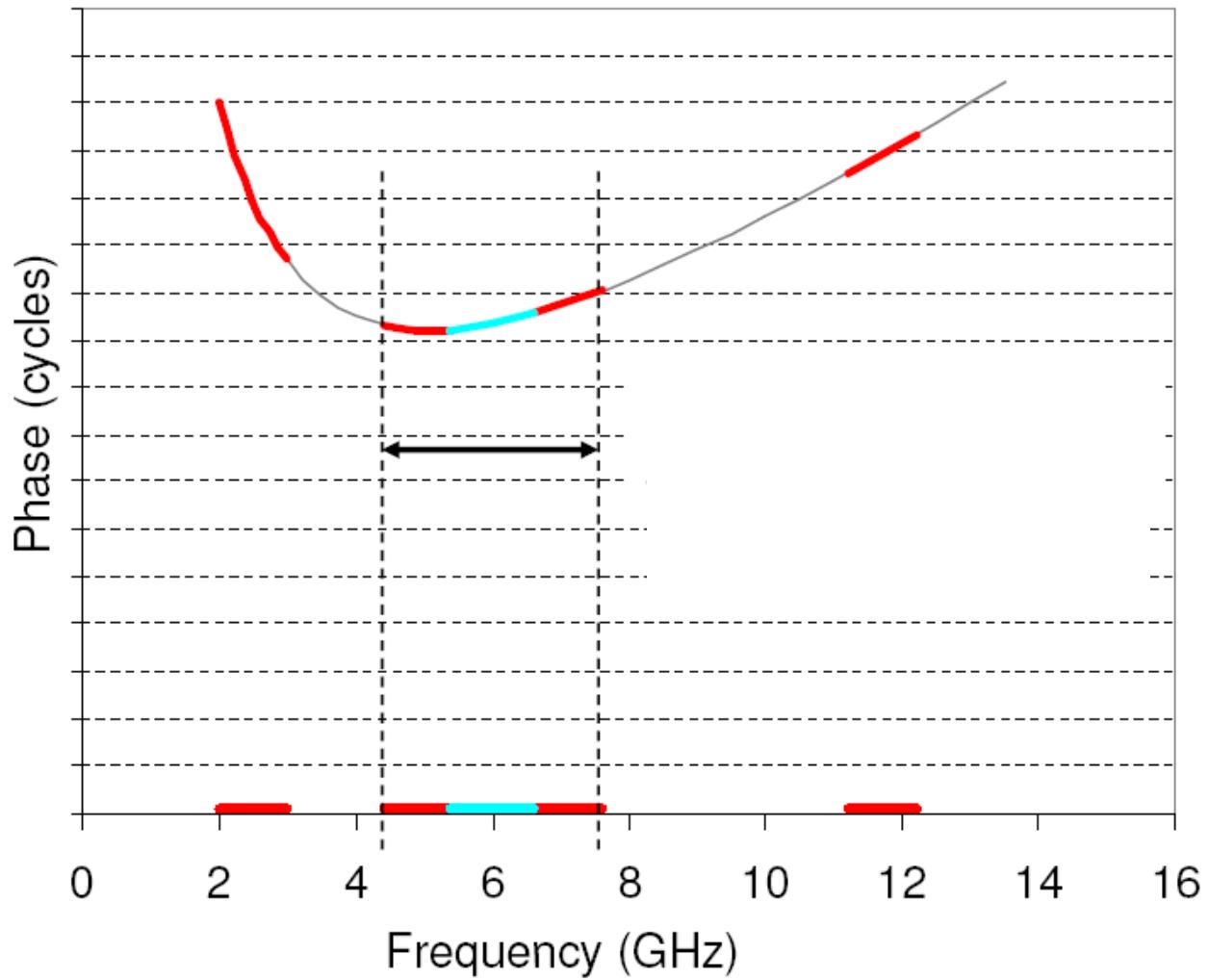


VLBI2010 Global Observing System
became
VLBI Global Observing System – VGOS

- Small and agile telescopes
- Large bandwidth (2 – 14 GHz)
- Flexible frequency allocation
- Dual linear polarization

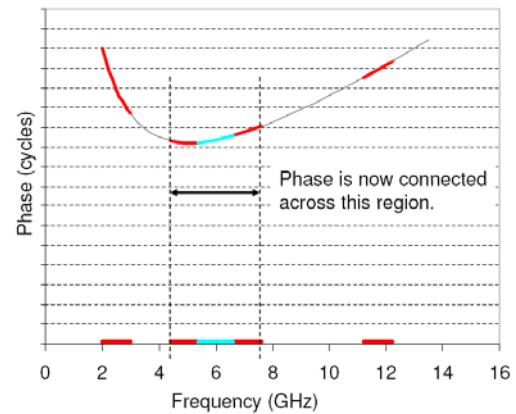


Frequency allocation for VGOS

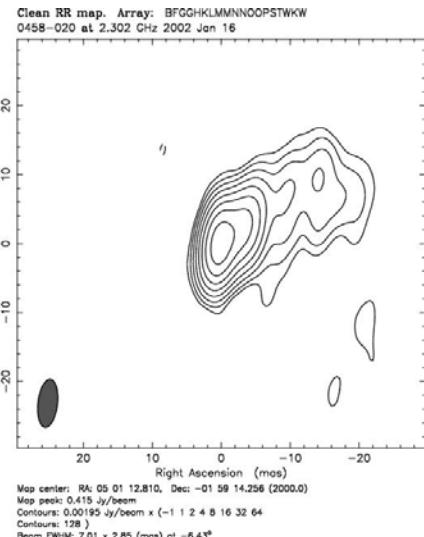


Challenges

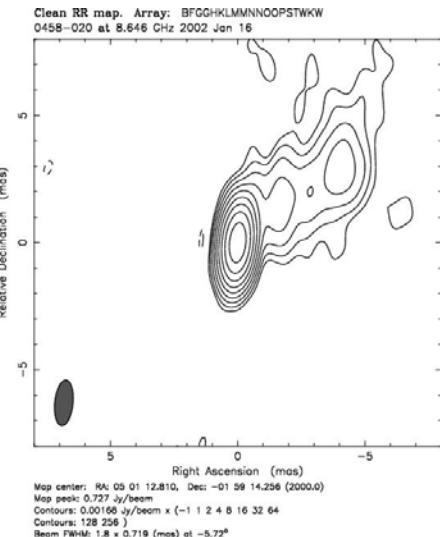
- Frequency band selection
 - Radio frequency interference
 - Phase connection requirements
- Source structure effects



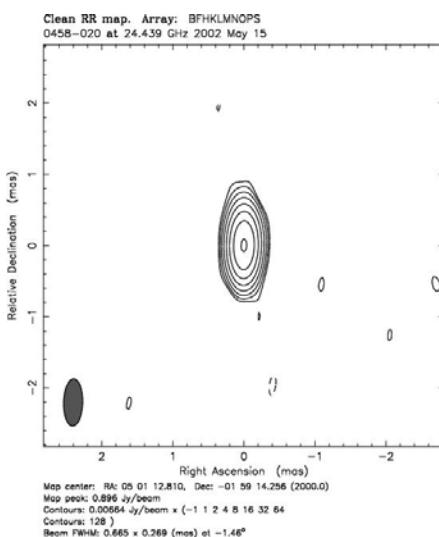
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S-band
2.3 GHz
13.6 cm



X-band
8.6 GHz
3.6 cm



K-band
24 GHz
1.2 cm

Images credit:
P. Charlot et al, AJ, 139, 5,
2010

New VGOS radio telescopes



NyAlesund (NO)

Courtesy L. Langkaas



Zelenchukskaya (RU)

Courtesy
A. Ipatov



Badary (RU)

Courtesy
A. Ipatov



Ishioka (JP) Courtesy Y. Fukuzaki

GGAO (US)

Courtesy A. Niell



STATUS OF THE SPANISH/PORTUGUESE RAEGE PROJECT



GOB^EERNO
DE ESPAÑA
MINISTERIO DE
EDUCACIÓN,
CIENCIAS Y
CULTURA



universität bonn



Gómez-González
et al .(2013)



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Yebes (Spain)
(August 2013) Courtesy: J.A. Lopez

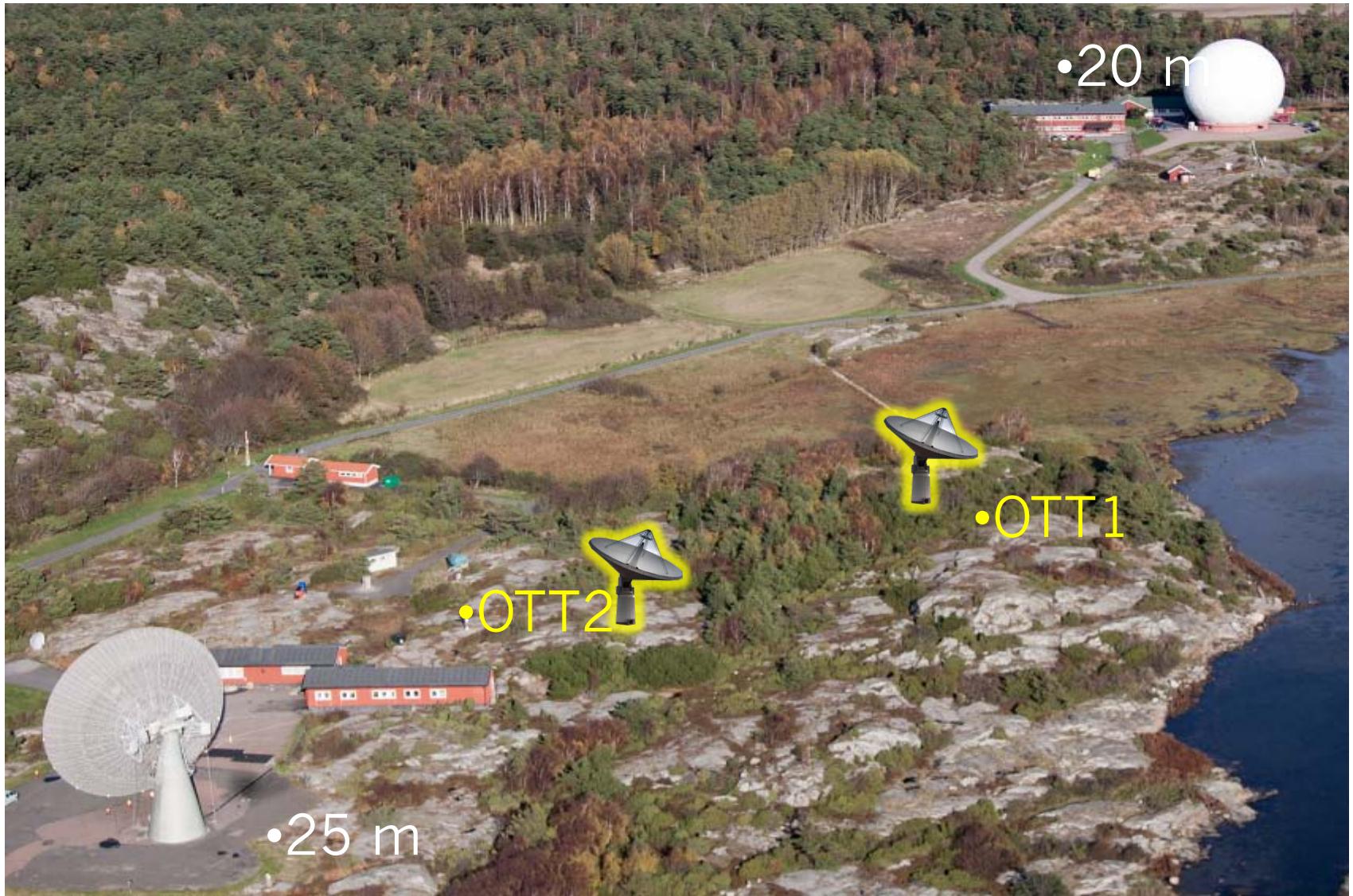


Santa Maria (Eastern Azores)
(Sep. 2014) Courtesy: F. Colomer

Twin Telescope Wettzell

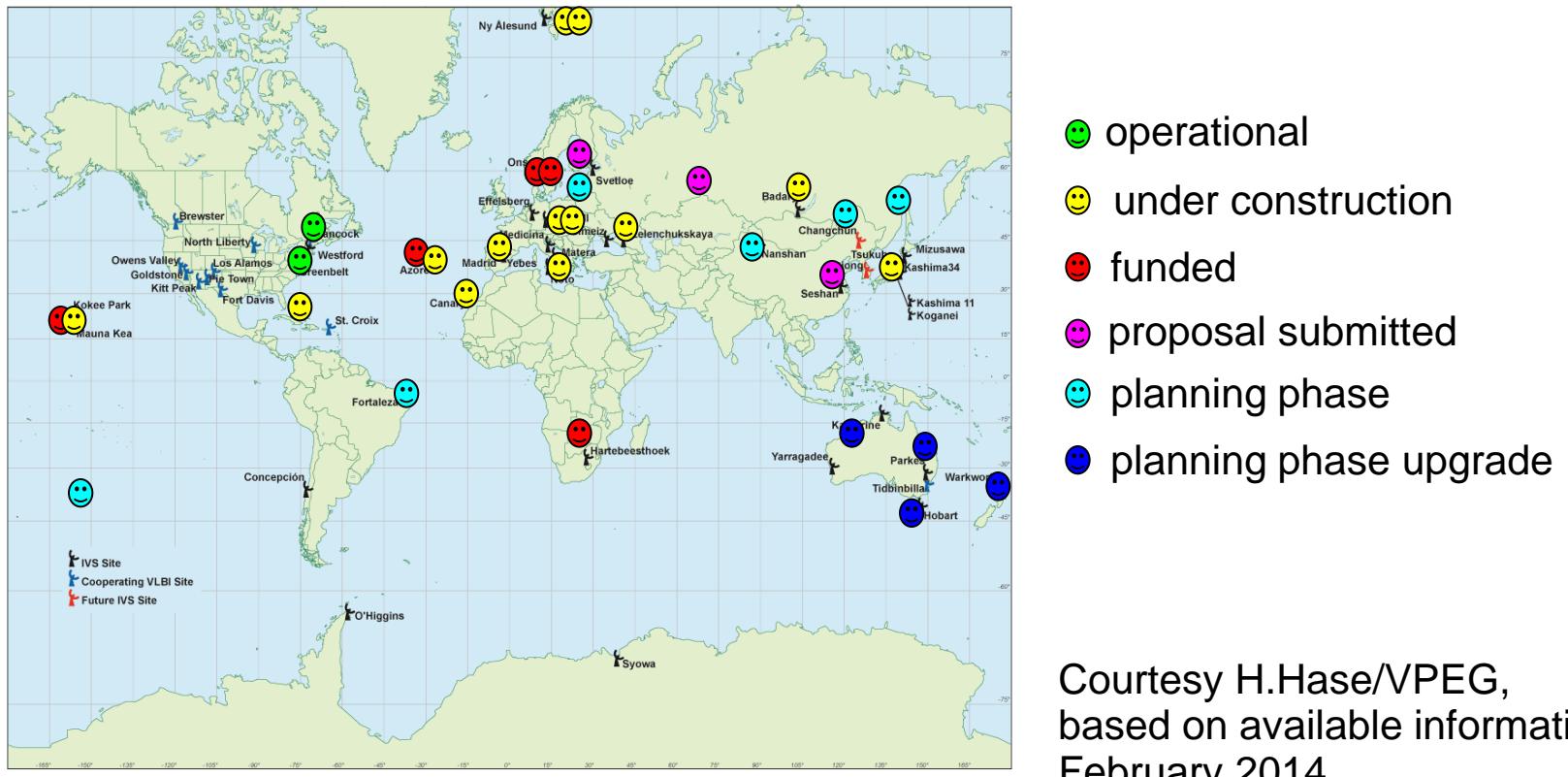


Twin Telescope Onsala



VGOS World

New VGOS radio telescopes for IVS



VGOS Observing Plan (Pilot Project)

VGOS Observing Plan (Densification Phase)

VGOS Observing Plan (Densification Phase)

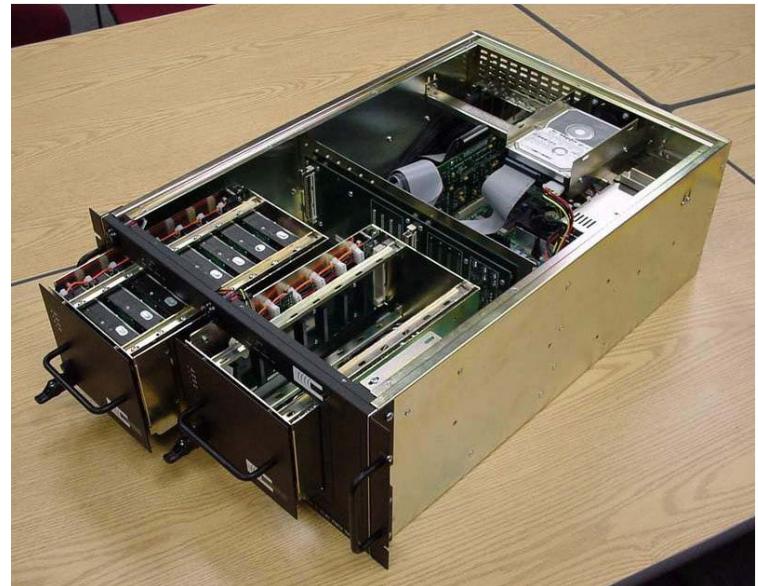
VGOS Observing Plan (Densification Phase)

VGOS Implementation Plan (Densification Phase)

VGOS Observing Plan (Densification Phase)

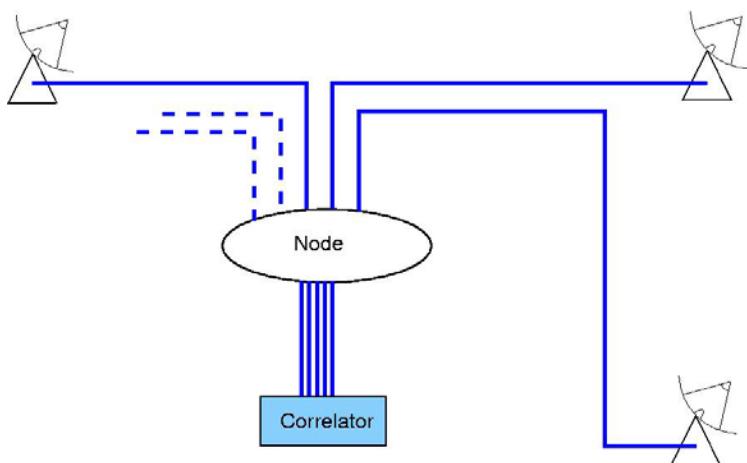
Options

- Shipping disk modules
- Electronic transfer



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Courtesy A. Whitney



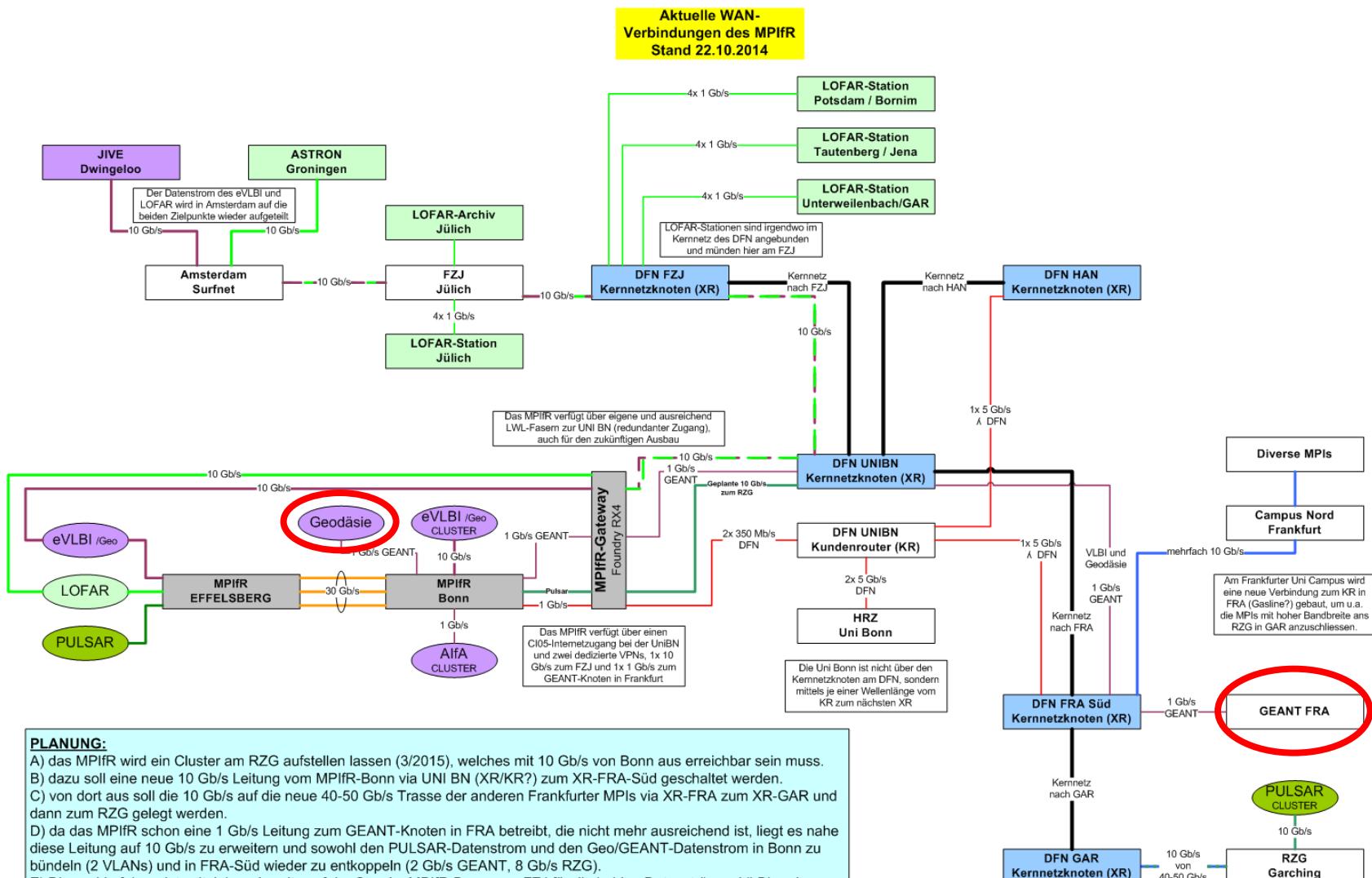
Correlator connectivity

- MIT Haystack (US) 20 Gb/s
- GSFC Tsukuba (JP) 10 Gb/s
- NEOS Washington DC (US), 1 Gb/s
- MPIfR/BKG Bonn (DE), 1 Gb/s

Year	# of sites	hours of obs/day	data/day/site (TB)	data/day at correlator (TB)
2015	8	4	7.2	58
2016	10	8	14.4	144
2017	16	8	14.4	230
2018	20	10	18.0	360
2019	24	12	21.6	518
2020	24	24	43.2	1037

Year	data rate at each site (Gbps)	network data rate at each site (Gbps)	data rate at correlator (Gbps)	network data rate at correlator (Gbps)
2015	0.7	1.0	5	8
2016	1.3	1.9	13	19
2017	1.3	1.9	21	30
2018	1.7	2.4	34	48
2019	2.0	2.8	48	68
2020	4.0	5.6	96	134

Situation at Bonn



Software Correlators



Estimated correlator cores required

Year	# of correlator cores
2015	200
2016	600
2017	900
2018	1400
2019	2000
2020	3900

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Expected developments

Location	Correlator Cores		External Network (Gbps)	
	Now	Planned	Now	Future
Bonn	488	1000-1500	1	??
USNO	512	1024	1	10
Haystack	100	~300	20	no plan
Shanghai	64	1000	1	no plan
Tsukuba	92	256	10	no plan

PC cluster with off-the-shelf components (scalable)

Challenge: Power consumption (for processors **and** for cooling)

Data analysis requirements in VGOS era

- Tremendous increase in observables
 - High degree of automatization required
 - Different levels of latency
 - Near real-time
 - Rapid
 - Final
 - Dependency on rapid availability of auxiliary data
 - Meteorological data
 - Mapping functions from numerical weather models
- IVS Task force on seamless auxiliary data archives

- VGOS Observing Plan
Petrachenko W, Behrend D, Hase H, Ma C, Niell A, Nothnagel A, Zhang X (adopted by IVS DB)
- VGOS Data Transmission and Correlation Plan 2014
Petrachenko W, Bertarini A, Alef W, Behrend D, Cappallo R, Hase H, Ma C, Niell AE, Nothnagel A, Zhang X (adopted by IVS DB)
- VGOS Analysis Plan
Gipson J et al. (in preparation)

See

<http://ivscc.gsfc.nasa.gov/technology/vlbi2010-documents.html>



International Association of Geodesy

Global Geodetic Observing System

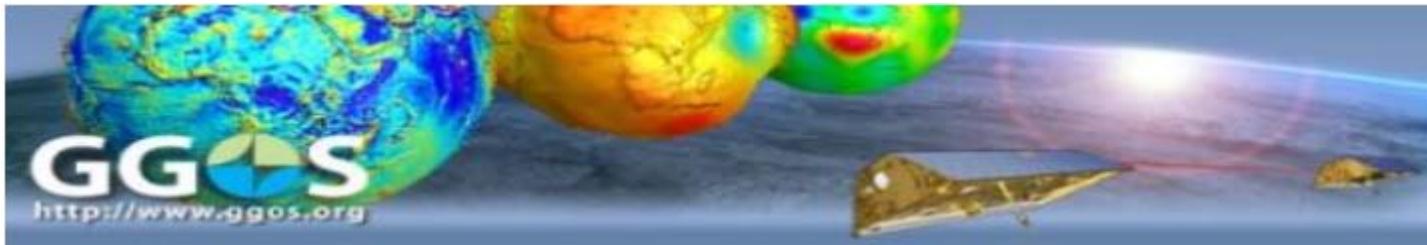


The mission of GGOS is:

- to provide the observations needed to monitor, map and understand changes in the Earth's shape, rotation and mass distribution;
- to provide the global frame of reference that is the fundamental backbone for measuring and consistently interpreting key global change processes and for many other scientific and societal applications;
- to benefit science and society by providing the foundation upon which advances in Earth and planetary system science and applications are built.



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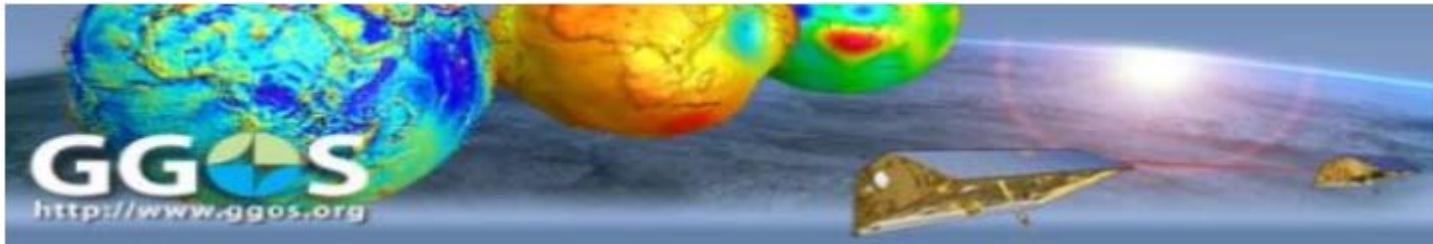


IVS → VGOS → GGOS

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Requirements of GGOS

- Global distribution
- Continuous
- Stable over decades
- 1 mm/0.1 mm/y



IVS → VGOS → GGOS

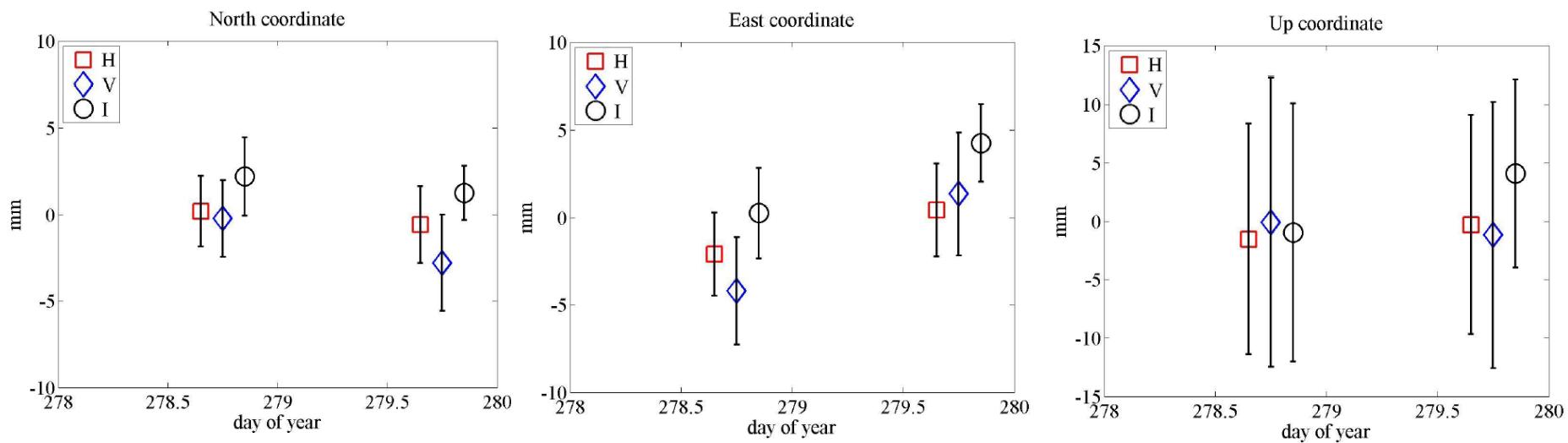
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Contribution to GGOS

- Global distribution → Well-designed network
- Continuous → Economic operations
- Stable over decades → Monitoring of telescopes and local ties
- 1 mm/0.1 mm/y → Improved technology, better modeling

Westford – GGAO baseline (600 km)
2x 6 hours, slightly different setups
8 Gbit/s with 32 standard disks (4 * Mark5C)

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The future has started

